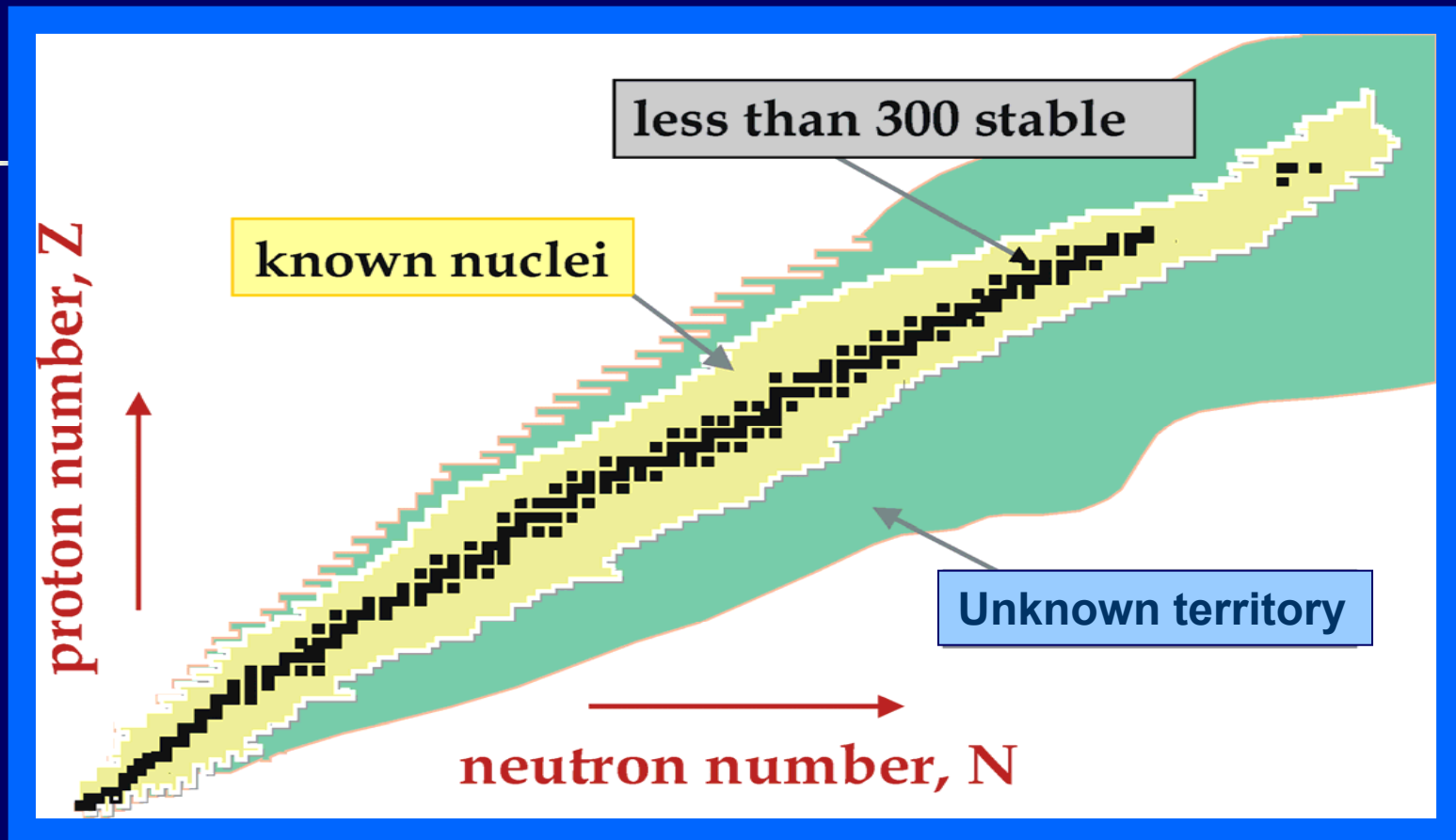


# Radioactive Ion Beam (RIB) Facility at VECC :

## Present & Future

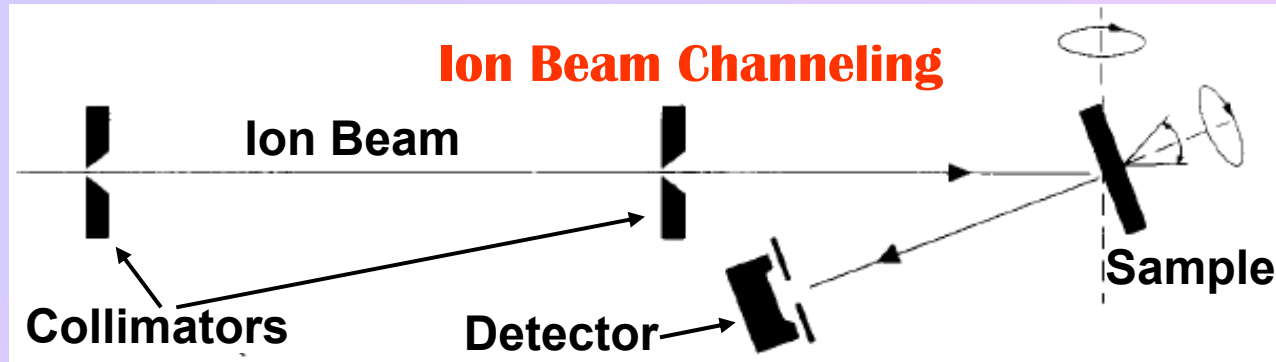
Arup Bandyopadhyay  
Variable Energy Cyclotron Centre  
Kolkata

## Radioactive Ion Beams : Ions of $\beta$ -unstable nuclei

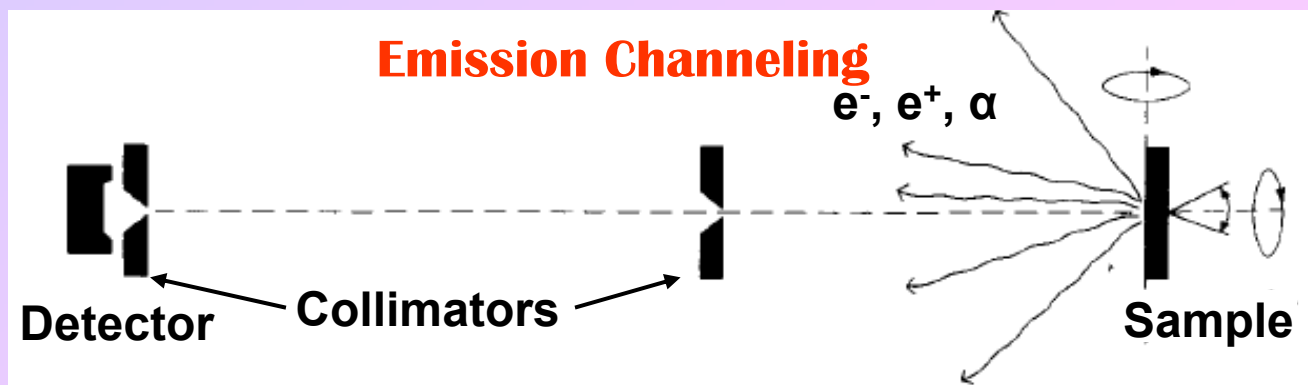


**Enormous increase in the no. of available projectiles**

## Study of formation & propagation of lattice defects



Well collimated ion beam interacts with lattice & impurity atoms – backscattered yield is measured as a function of sample orientation



Yield of charged particles emitted by radioactive impurity atoms is measured as a function of sample orientation

### Advantages of EMISSION CHANNELING technique :

Req. implantation dose of radioactive atoms is significantly lower than that of ion channeling experiment.

➤ Radiation damage during channeling analysis negligible.

➤ **Sensitivity of EC >> higher than ion channeling tech. ( $1E13$  &  $1E18$  /cc)**

## Physics Motivation

- **Material Science**

**Mössbauer Effect (ME)** : Study of recoil-free emission & absorption of  $\gamma$  by a radioactive nucleus embedded in a solid.

**ME allows measurement of Isomer Shift (IS)** - relative shift of the emission and absorption lines proportional to the difference in electron density at the Mössbauer nucleus in the source and in the absorber matrix

→ Provides information about the lattice site of the Mössbauer atom & its charge state.

### Advantage of using RIB as a Mössbauer dopant are :

➤ **Chemical incompatibility of Mössbauer atom is not an issue any more : possible to find a suitable compatible parent isotope which decays to the Mössbauer atom.**

➤ **Site selective doping which depends on the parent isotope**

**For example,  $^{119}\text{Sn}$  – good Mössbauer dopant for group III-V compound semiconductors :**

**$^{119}\text{In} \rightarrow ^{119}\text{Sn}$  (occupies group V lattice sites)**

**$^{119}\text{Sb} \rightarrow ^{119}\text{Sn}$  (occupies group III lattice sites)**

**Radioisotope Therapy** : A radionuclide is delivered to a tumor site using a biologically active molecule which decays via  $\beta^-$  or  $\alpha$ .

Using RIB one can produce such nuclide with **high specific activity** as it will be carrier free or no-carrier-added form (suitable target+projectile & clean separation) + **availability of new radioisotopes**.

**PET (Positron Emission Tomography)** – medical imaging of tumors, mapping of human brain and heart function. Most of the PET isotopes are short lived for clinical use & research. Using RIB it is possible to produce **longer lived PET isotopes** ( $^{72}\text{As}$  :  $T_{1/2} \sim 26$  hours) which are carrier free i.e. **high sp. Activity**.

# Physics Motivation

- **Nuclear Physics**

## Study of nuclei away from line of $\beta$ -stability

- **Neutron Halo** ( ${}^6,{}^{11}\text{Li}$ ,  ${}^{14}\text{Be}$ ,  ${}^{17}\text{B}$ , .....
- **$\beta$ -delayed particle emission** ( ${}^{21}\text{Mg}$ ,  ${}^{16}\text{C}$ ,  ${}^{22}\text{Al}$ ,  ${}^{26}\text{P}$ , ...)
- **Nucl. Deformation & magicity** (  ${}_{40}^{80}\text{Zr}$ ,  ${}_{38}^{100}\text{Sr}$  )

- **Nuclear Astrophysics**

- **The nuclear processes responsible for nucleosynthesis**
- **When & where these processes took place**
- **Characteristics (temperature, density, composition ..) of the nucleosynthesis sites**



## Primordial Nucleosynthesis

The Universe was created about 15 billion years back by gradual expansion from a state of extremely high density ( $\rho \sim 10^{95}$  g/cc) & temperature ( $T \sim 10^{32}$  K) → Big Bang Theory.

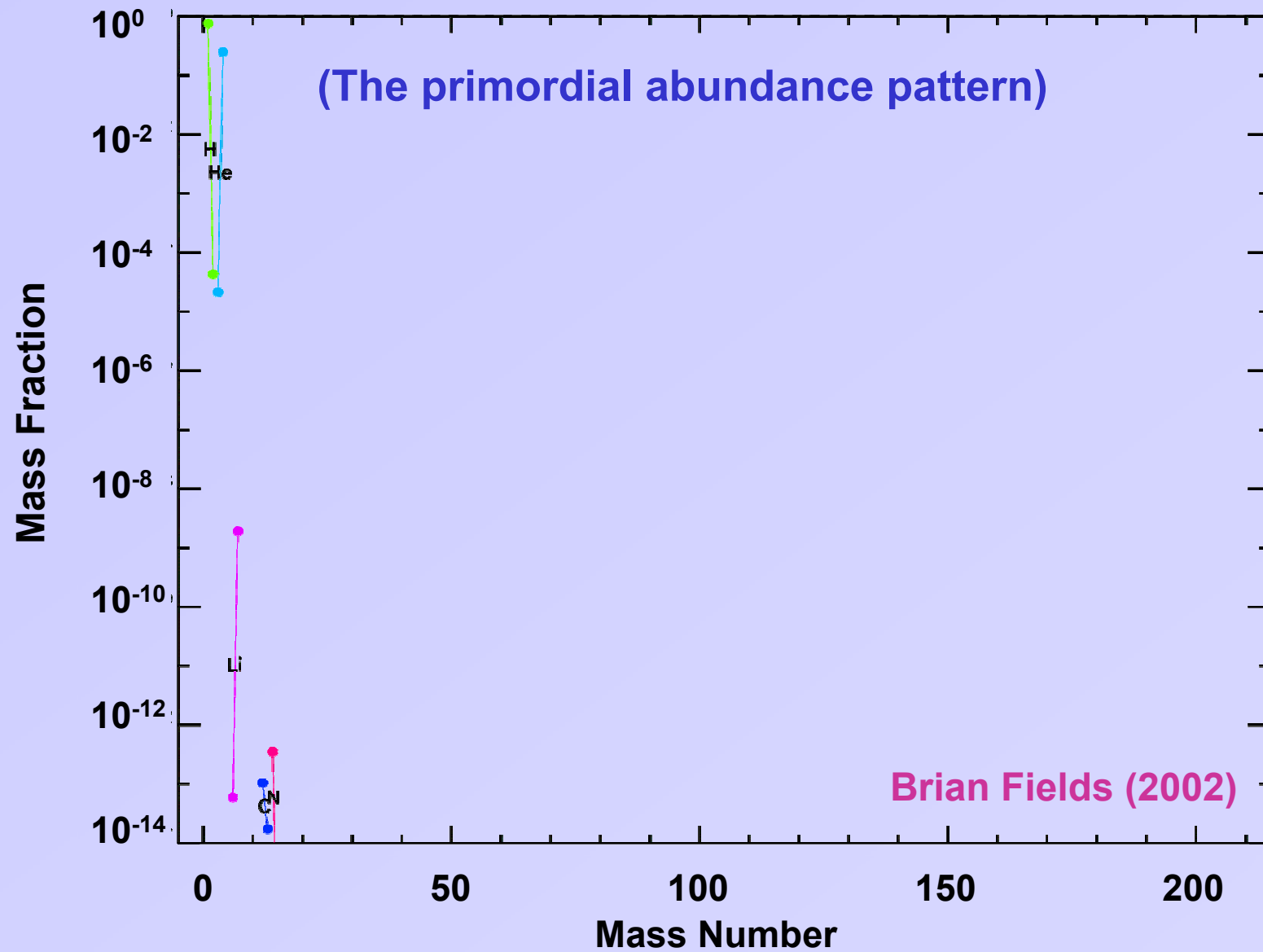
Nucleosynthesis could start ~ 3 minutes after the Big Bang when Deuterium (B.E. 2.23 MeV) stable against photo-dissociation at  $T \sim 10^9$  K.

Nucleosynthesis could continue for a very short time after the big-bang because :

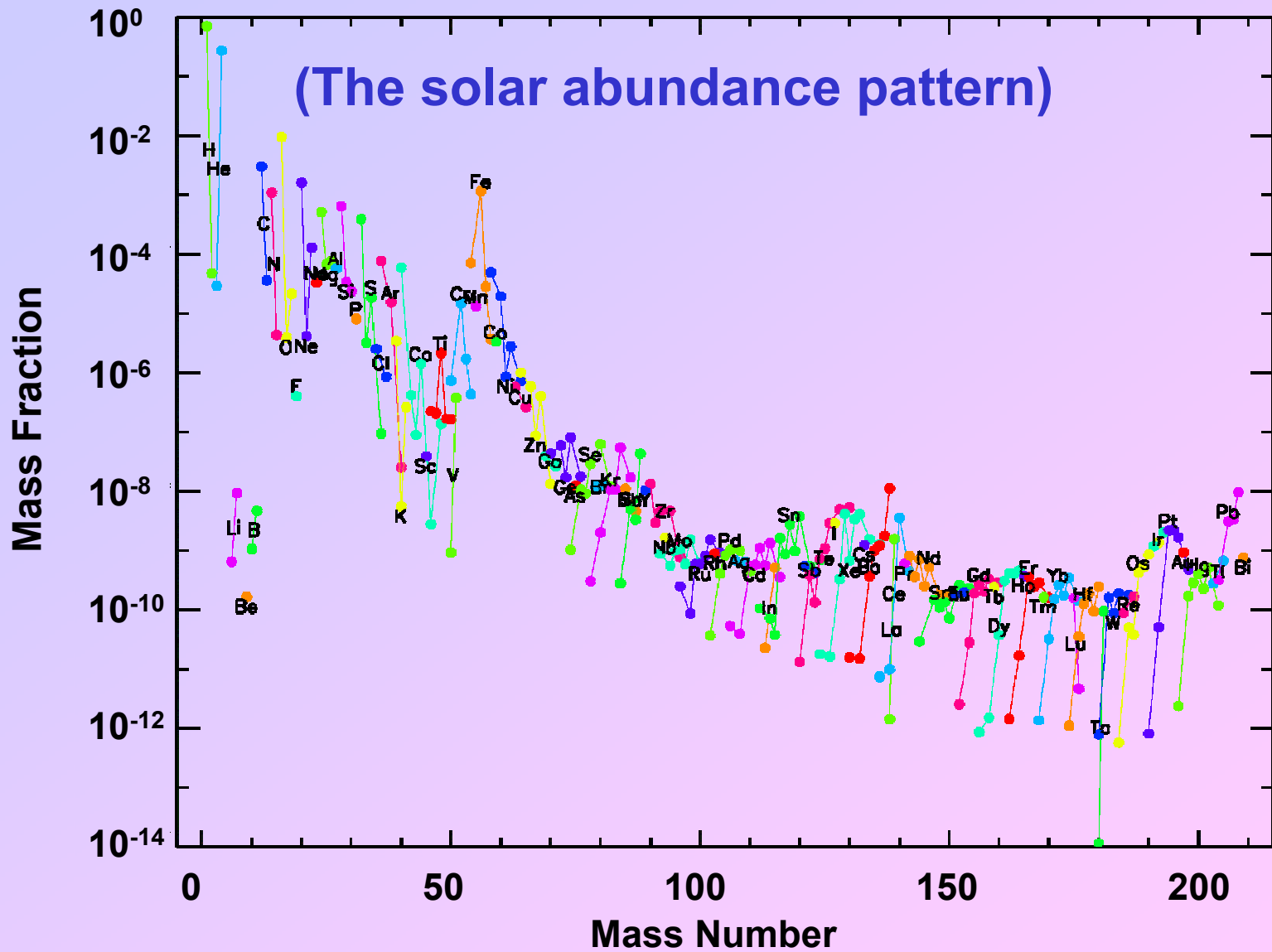
- All neutrons were used up for production of  ${}^4\text{He}$
- No stable nuclei for  $A=5$  and  $A=8$
- Baryon density too low for 3 body reactions
- Temp. reduced due to expansion to sustain further nuclear reaction

Mass Fraction	
${}^1\text{H}$	0.75
${}^2\text{H}$	$(2.5 \pm 1.5) \times 10^{-5}$
${}^3\text{He}$	$(4.2 \pm 2.8) \times 10^{-5}$
${}^4\text{He}$	$0.23 \pm 0.02$
${}^6\text{Li}$	$300^{+300}_{-150} \times 10^{-12}$
${}^7\text{Li}$	$4600^{+4600}_{-2300} \times 10^{-12}$

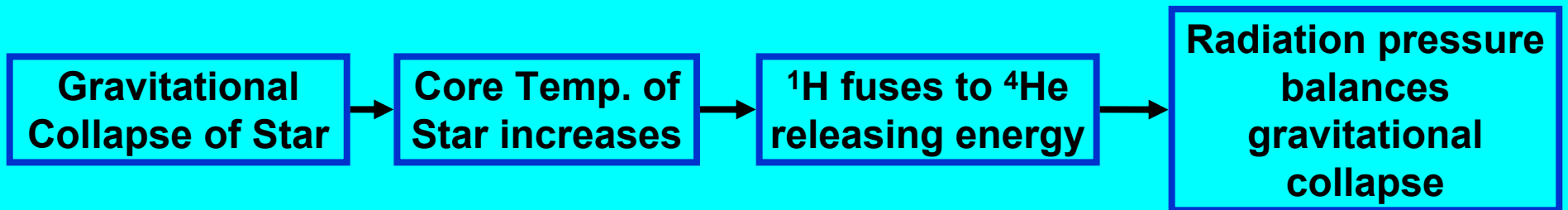
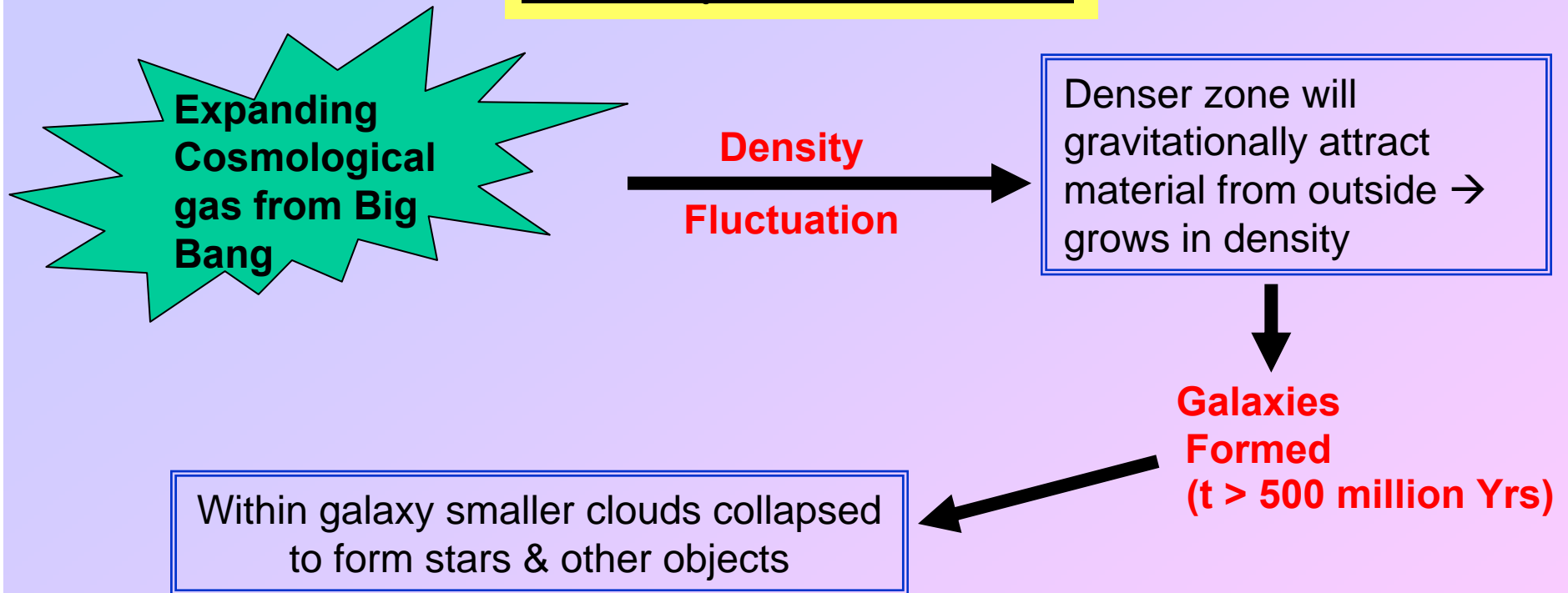
# Primordial Nucleosynthesis



# Nucleosynthesis



## Nucleosynthesis in Stars



**Hydrogen burning : Source of energy of Main Sequence Stars like Sun – continues for tens of millions of years**

# Nucleosynthesis in Stars

Elements are created in several stages where the products of one stage become the fuel for energy generation in the next stage.

$$M < 8M_{\odot}$$

After He burning stage, star will throw off its envelop as a planetary nebula & become a white dwarf.

$$M > 8M_{\odot}$$

Main sequence → Goes through all burning stages → Red Giant → Core of mostly Fe → gravitational collapse to super-density → Supernova explosion → Blows off outer envelop → remnant as neutron star / black hole

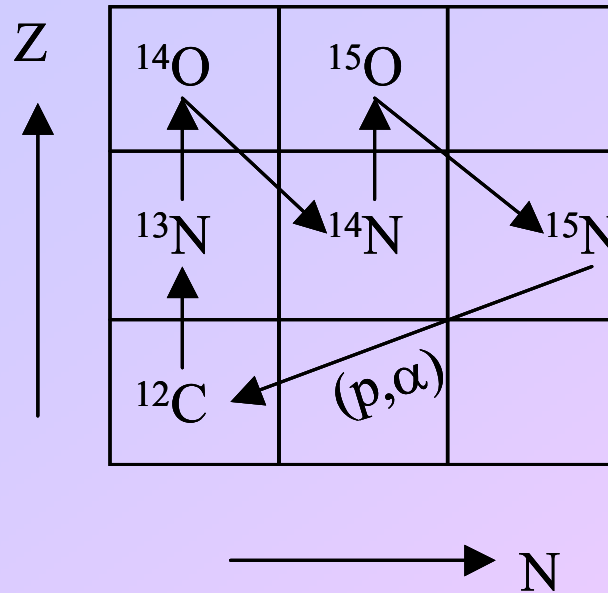
## Hydrostatic burning stages

<i>H burning</i>	$4p \Rightarrow {}^4\text{He} + 2e^+ + 2\nu + 26.7 \text{ MeV}$
<i>He burning</i>	$\alpha + \alpha \Rightarrow {}^8\text{Be}$ ${}^8\text{Be} + \alpha \Rightarrow {}^{12}\text{C} + \gamma + 7.2 \text{ MeV}$ ${}^{12}\text{C} + {}^{12}\text{C} \Rightarrow {}^{23}\text{Na} + p + 2.23 \text{ MeV}$ $\Rightarrow {}^{20}\text{Ne} + \alpha + 4.6 \text{ MeV}$
<i>C burning</i>	${}^{12}\text{C}(p, \gamma){}^{13}\text{N}(e^+ \nu){}^{13}\text{C}(\alpha, n){}^{16}\text{O}$
<i>Ne burning</i>	${}^{20}\text{Ne} + \gamma \Rightarrow {}^{16}\text{O} + \alpha$ ${}^{20}\text{Ne} + \alpha \Rightarrow {}^{24}\text{Mg} + \gamma$
<i>O burning</i>	${}^{16}\text{O} + {}^{16}\text{O} \Rightarrow {}^{31}\text{P} + p$ $\Rightarrow {}^{28}\text{Si} + \alpha$ $\Rightarrow {}^{31}\text{S} + n$
<i>Si burning</i>	${}^{28}\text{Si} + \alpha \Rightarrow {}^{32}\text{S} + \alpha$ $\dots + \alpha \Rightarrow {}^{56}\text{Fe}$ ${}^{28}\text{Si} + \gamma \Rightarrow \alpha + \text{Mg}$

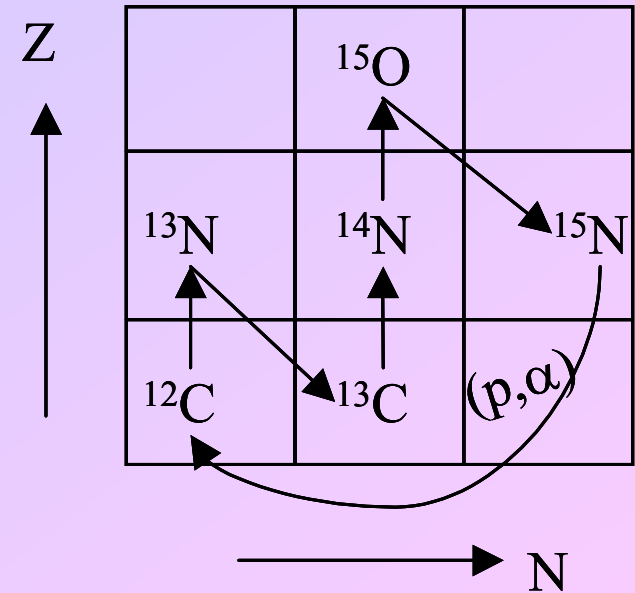


# Nucleosynthesis in Stars

## CNO Cycle



HOT CNO

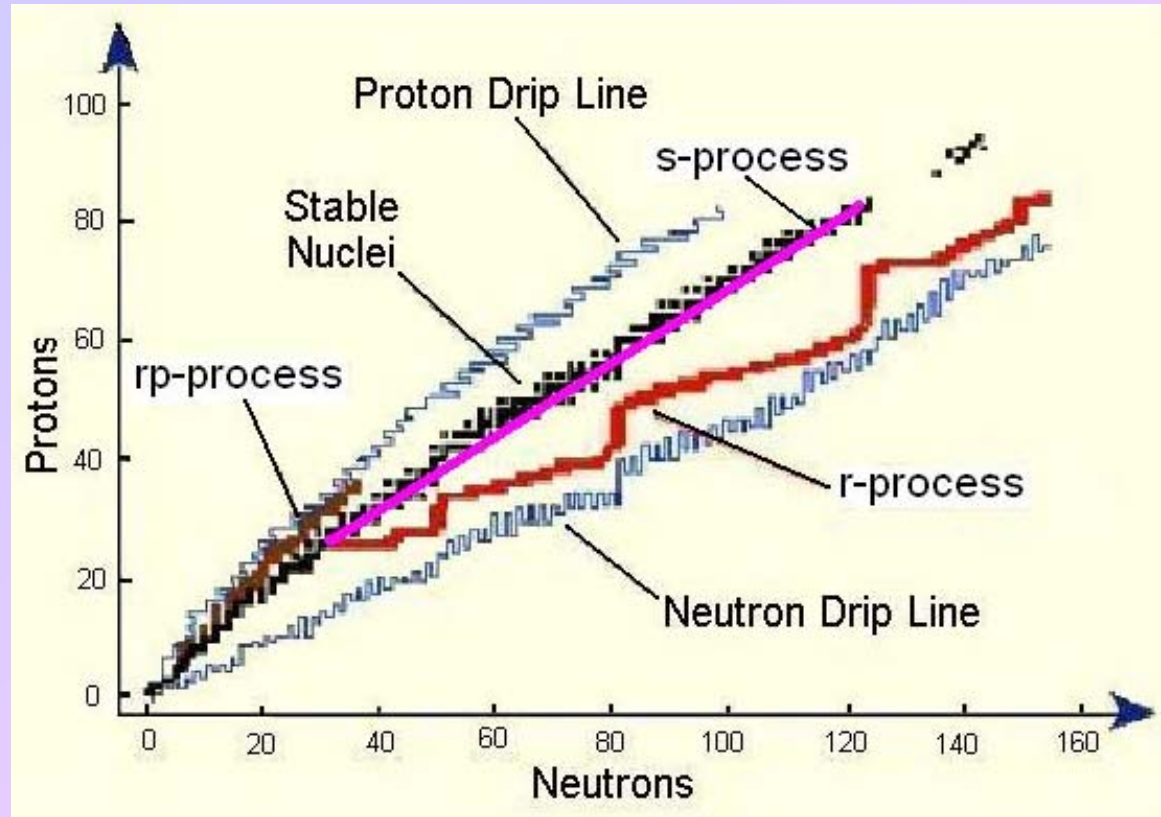


COLD CNO

A form of Hydrogen burning chain reaction takes place in non-first generation stars where C is available to start with.

## Nucleosynthesis of elements beyond Fe

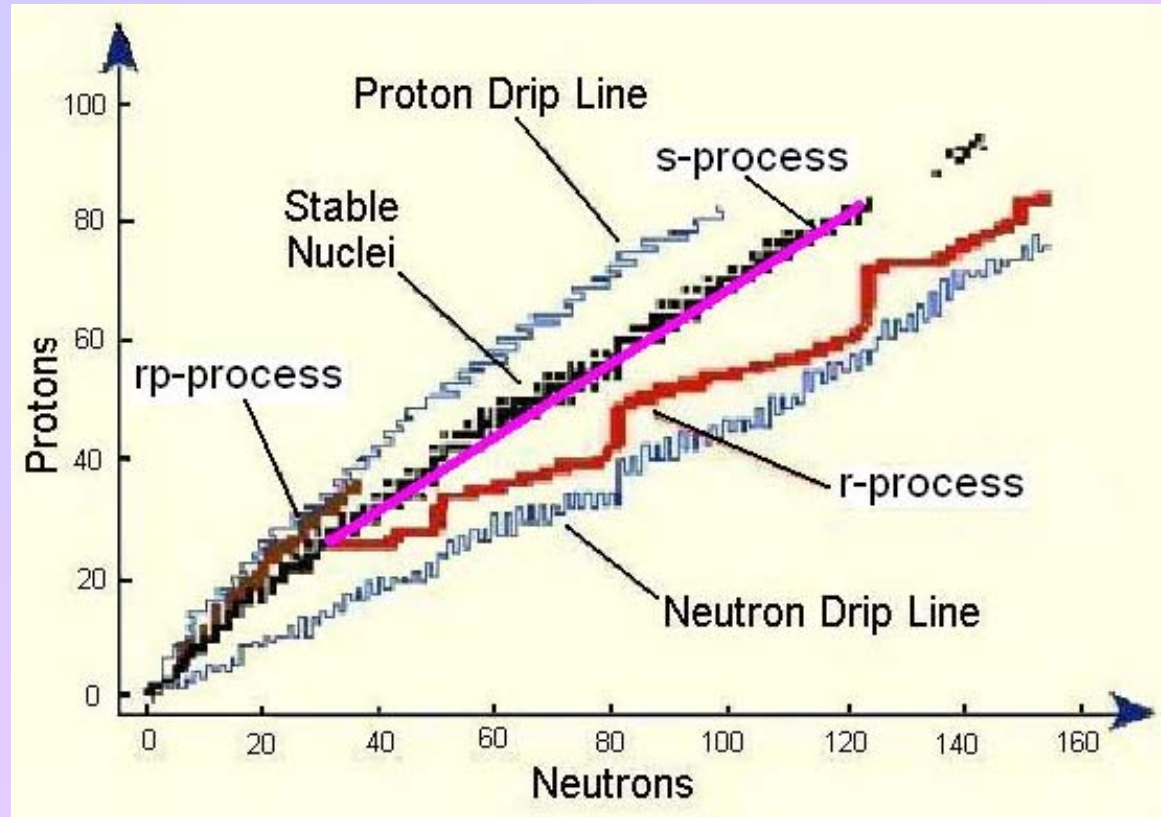
### r - Process



- Rapid successive n-capture ( $n, \gamma$ ) W/O intermediate  $\alpha$  and  $\beta^-$  decay
- Follows path closer to n-drip line
- Near magic nos. neutron BE is small –  $\beta^-$  decay occurs – path moves towards stability line
- Produces n-rich nuclei between  $A=56$  to 270 – beyond which fission sets in

## Nucleosynthesis of elements beyond Fe

### s - Process

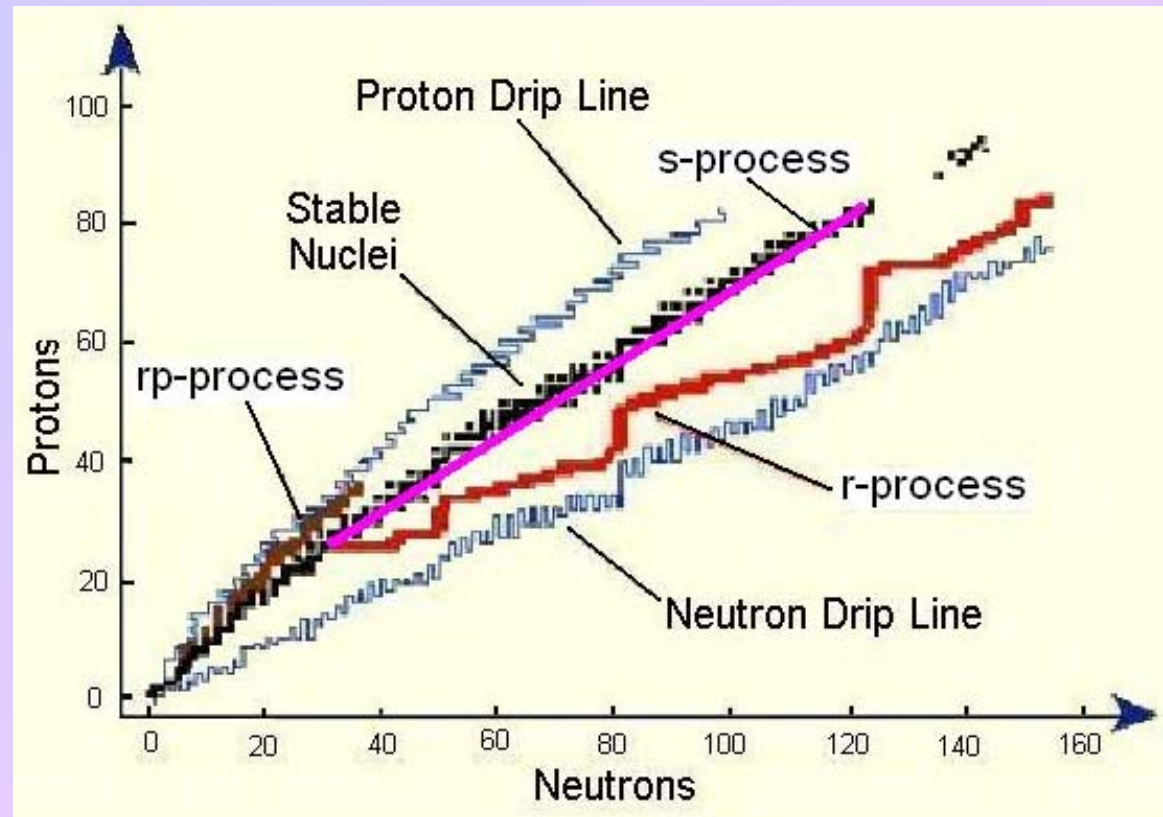


- Slow n-capture : slow w.r.t.  $\beta^-$  decay process
- Series of (n.g) processes and  $\beta^-$  decays (sometimes  $\beta^+$  and EC)
- Follows path closer to the stability line
- Produces n-rich nuclei between  $A=56$  to 209 (Bi) – beyond which there are many short lived nuclei & slow process cannot compete.
- Red giants – being a good source of neutrons from  $^{13}\text{C}(\alpha, n)^{16}\text{O}$ ,  $^{17}\text{O}(\alpha, n)^{20}\text{Ne}$ ,  $^{21}\text{Ne}(\alpha, n)^{24}\text{Mg}$  ... reactions– possible site for such processes.



# Nucleosynthesis in Stars

## rp - Process



- Rapid p-capture process : fast w.r.t.  $\beta^+$  decay process
- Follows path closer to the proton drip line
- Produces p-rich nuclei up to  $A \sim 100$  – beyond which Coulomb repulsion prevents successive p-capture.

# Nuclear Astrophysics

- Reactions involving stable nuclei drive the evolutionary stage of main sequence stars
- Reactions involving unstable nuclei dominate the outcome of explosive events

## Major Challenges

- Number of target atoms/cm<sup>2</sup> is small as reactions have to be studied at a very low energy, thin targets have to be used
- Expected cross section is small near or below the Coulomb barrier
- Intensity of RIBs is generally less compared to stable beams
- Often the RIBs are not pure (mixed beam) & have large energy spread
- Background due to the decay of RIB

# Nuclear Astrophysics

## Major Detection Techniques Needed

- Highly segmented high efficiency gamma ray detector array
- Large solid angle and good angular resolution silicon detector array
- Mass separator with high beam rejection efficiency
- Gas cell targets & targets of long lived radioactive nuclei

**GRETA (Gamma Ray Energy Tracking Array) - LBL**

Solid angle coverage : 0.45 → 0.8

Detection Efficiency : 10→50 %

Position resolution : 20mm → 2mm

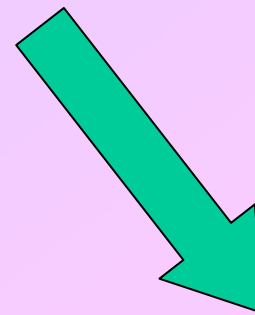
**DRAGON (Detector of Recoils And Gammas Of Nuclear reactions)**

Angular acceptance  $\leq \pm 20$  mrad

Energy acceptance  $\leq \pm 4\%$

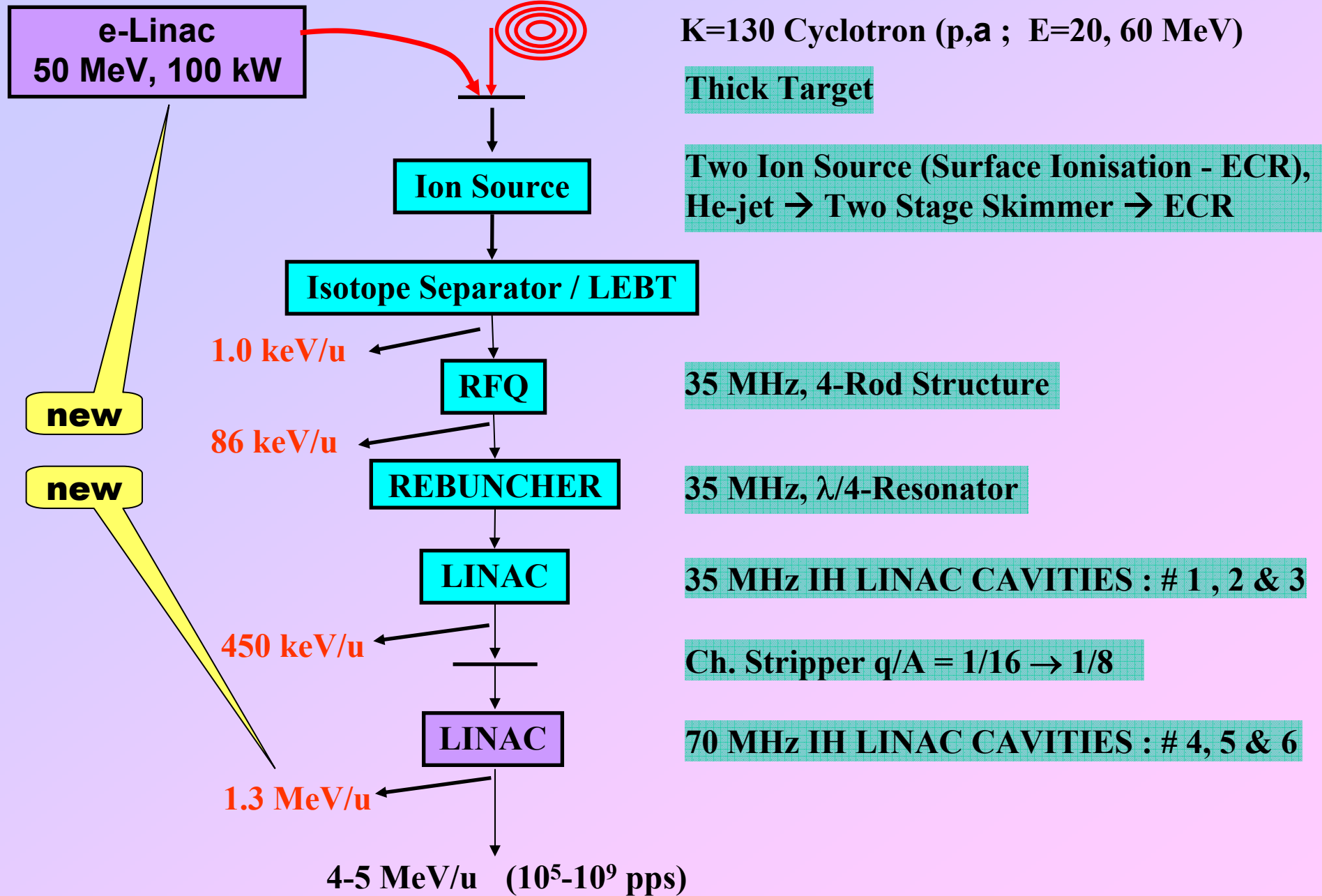
Separation is 1 in  $10^{15}$

1. Radioactive Ion Beams will play an important role in all fields of accelerator based research for the next few decades
2. Nuclear astrophysics is a fascinating field of research of tracing back the evolution of the Universe. It requires information of different reactions using both stable and radioactive beams.
3. But --- it requires stretching the limits of present accelerator and detection technologies.

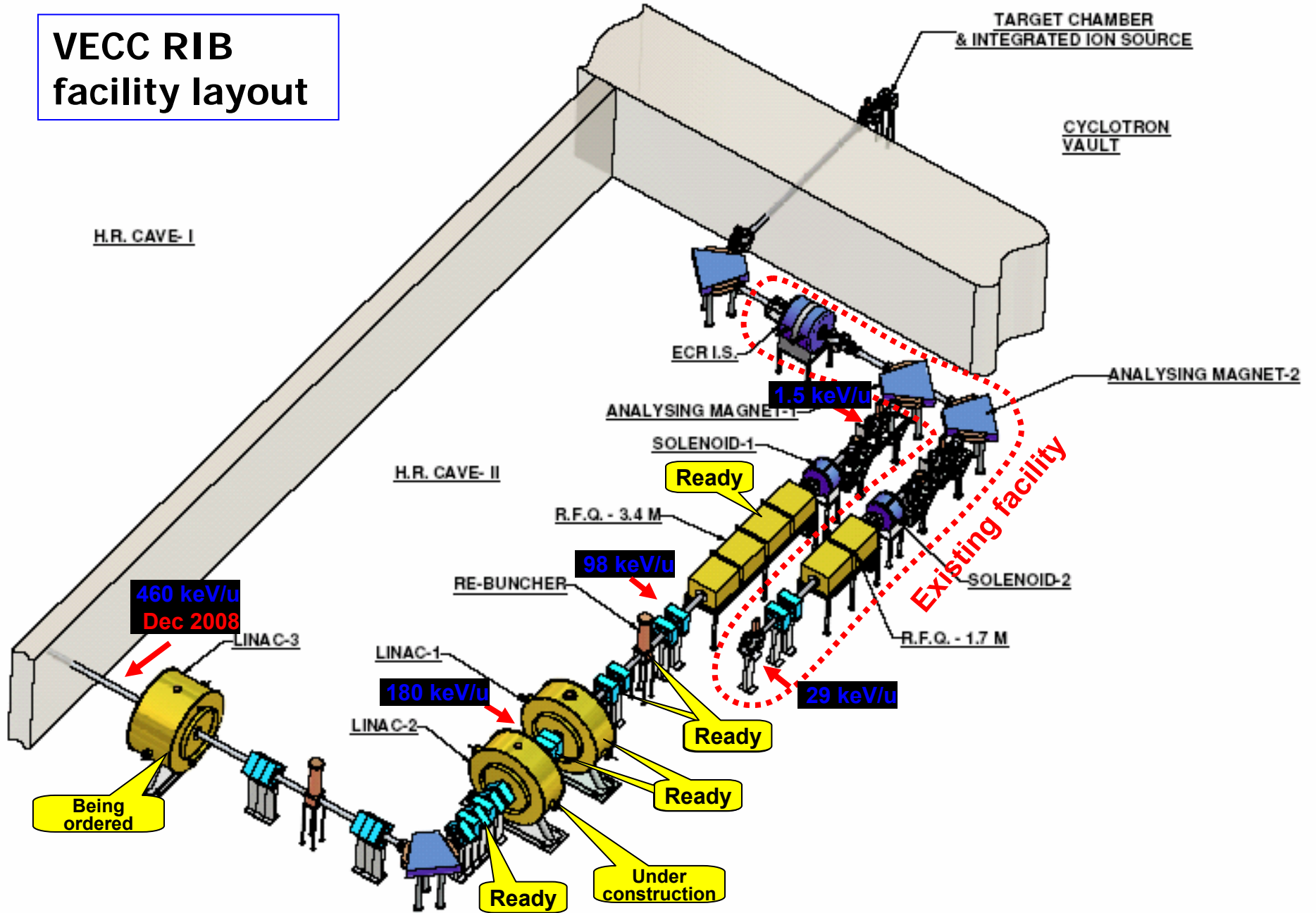


**How to produce RIB?**

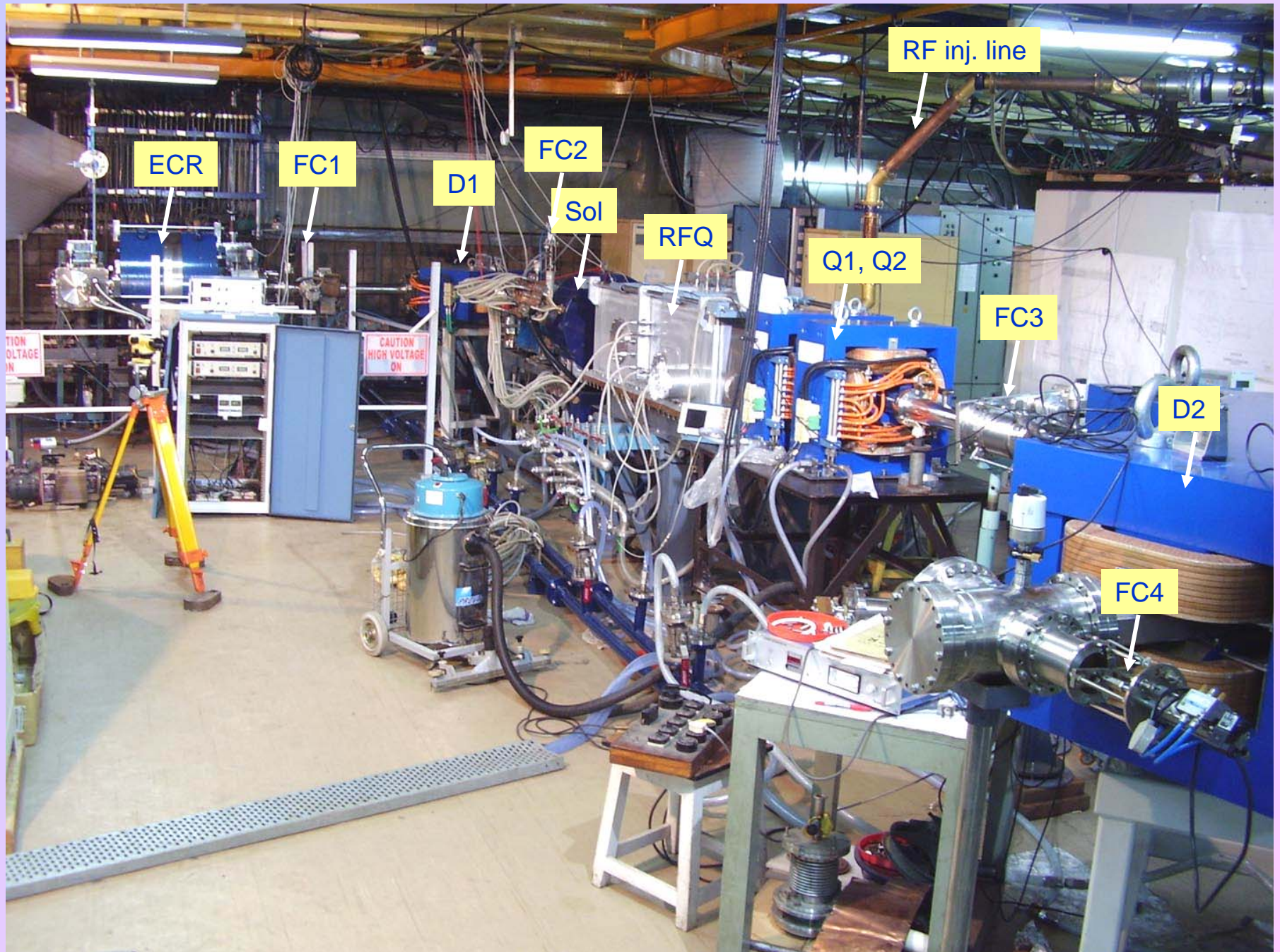
# Scheme of the RIB facility at VECC



# VECC RIB facility layout







## Thick target development :

### → Selection of suitable target material

ThermoCalc, Chemsage, HSC

(Data based thermo-chemistry & thermodynamic codes)

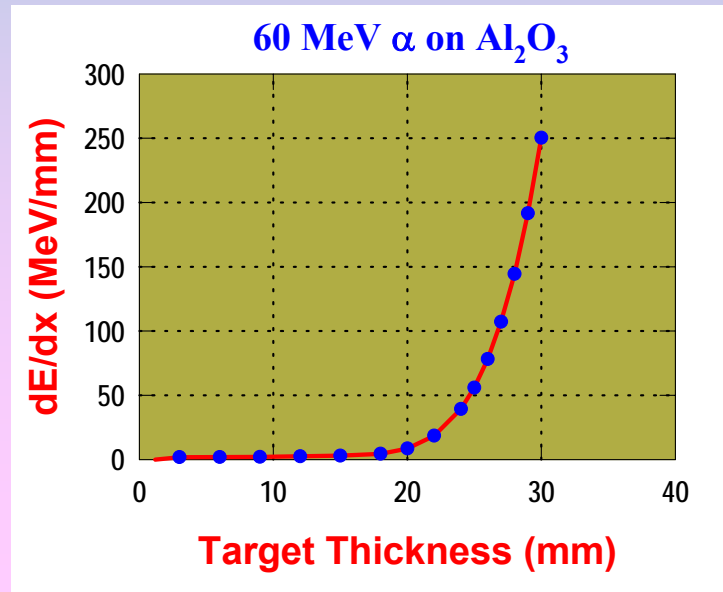
High dissociation temperature and low vapour pressure

### → Optimisation of the target geometry

Analysis of time dependent diffusion rate solving Fick's second equation for different symmetries like planar, cylindrical and spherical

### → Optimising target thickness

Heat deposition in the target is calculated using TRIM - heat deposition within the target sharply increases near the Bragg peak window.





# Thick target development :

→ Temperature distribution (ANSYS)

The critical temperature is determined from critical vapour pressure

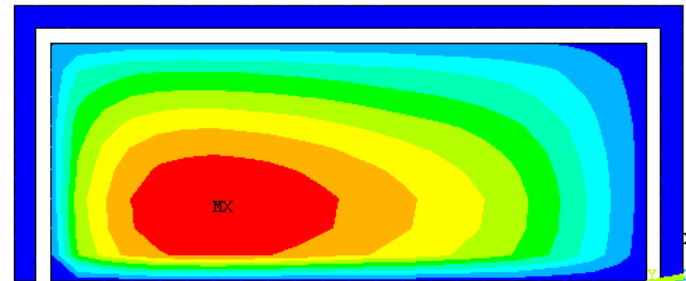
→ Maximise surface to volume ratio of the target i.e. porous / fiber like target matrix

1

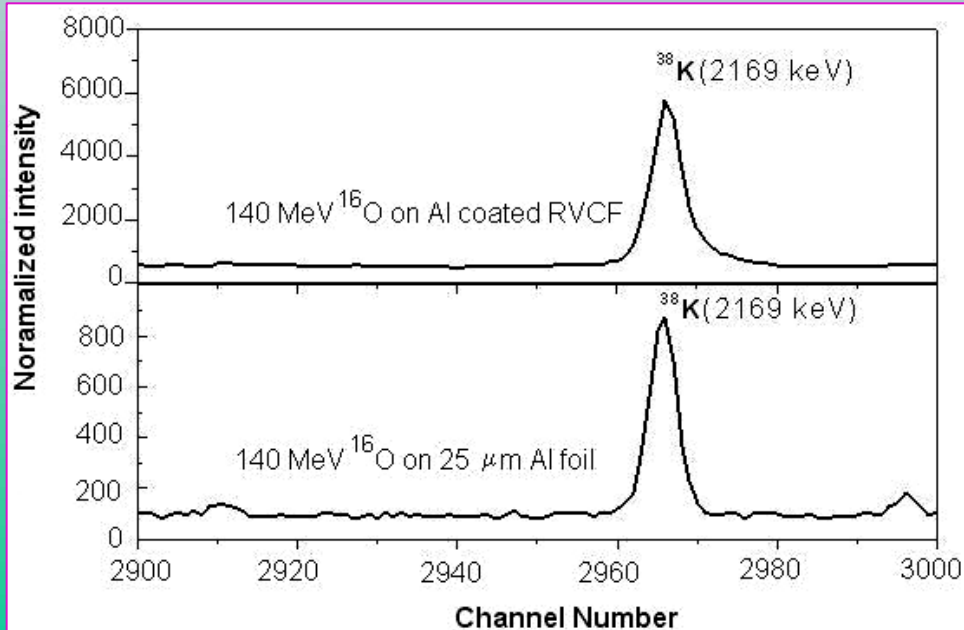
$\alpha \rightarrow \text{Al}_2\text{O}_3$  (  $T=2072^\circ \text{C}$  )

E/I : 60 MeV / 1 mA

0.75 cm f x 2cm Long

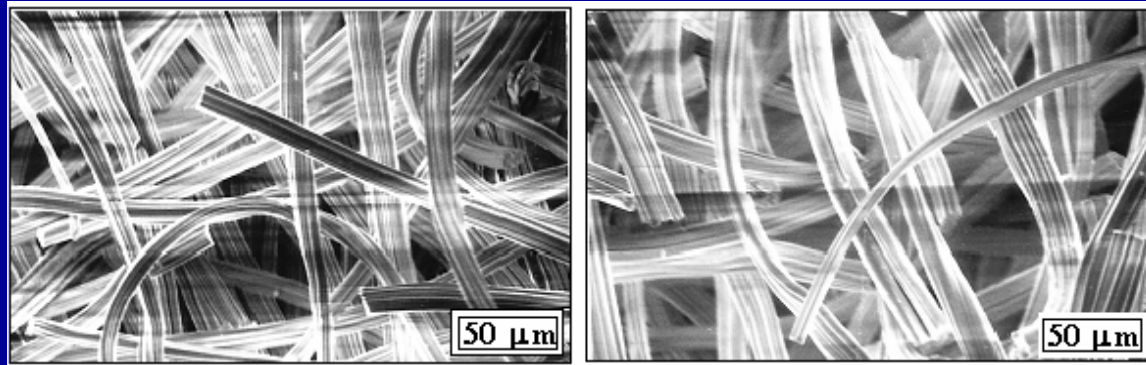


ANSYS 5.4  
FEB 7 2002  
13:24:46  
NODAL SOLUTION  
STEP=1  
SUB =1  
TIME=1  
TEMP  
TEPC=68.238  
SMN =100  
SMX =1392  
100  
243.568  
387.135  
530.703  
674.271  
817.839  
961.406  
1105  
1249  
1392

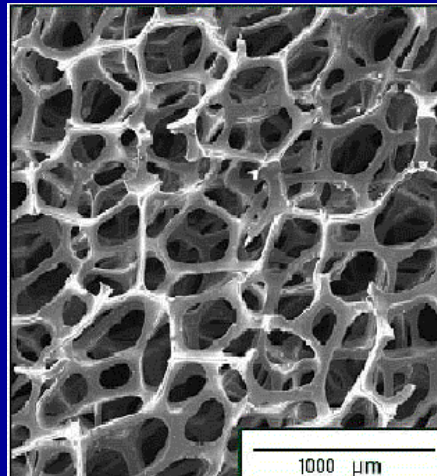


## Thick target R&D : first few targets

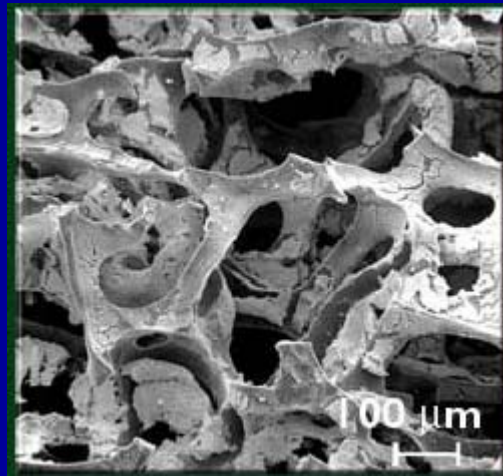
Carbon\* ,  $\text{Al}_2\text{O}_3$ , ZnO,  $\text{HfO}_2$ , BN, LiF, MgO,  $\text{CaCl}_2$ ,  $\text{ThC}_2$ ,  $\text{UC}_2$ ,  $\text{ZrO}_2$



SEM of  $\text{Al}_2\text{O}_3$  &  $\text{HfO}_2$



SEM of RVCF



SEM of RVCF +  $\text{Al}_2\text{O}_3$

\*RVCF : Reticulated Vitreous Carbon Fiber

## R&D Initiative on ion source :

- **Aim** → High On-line Efficiency for high charge States ( $q/A > 1/14$ )
- **Apparent Choice** → ECR ion source

**Problems**

- Poor vacuum (target evaporation) → high  $q$  !!!
- Neutron damage of permanent magnets

**Possible solution** → **Two ion source philosophy**

**Principle**

- Two I.S. in cascade – EBPIIS & ECR in our case
- First I.S. provides  $q=1^+$  even in poor vacuum
- Transported to ECR at a distance to get high  $q$  (Better vacuum and acceptable n-flux)

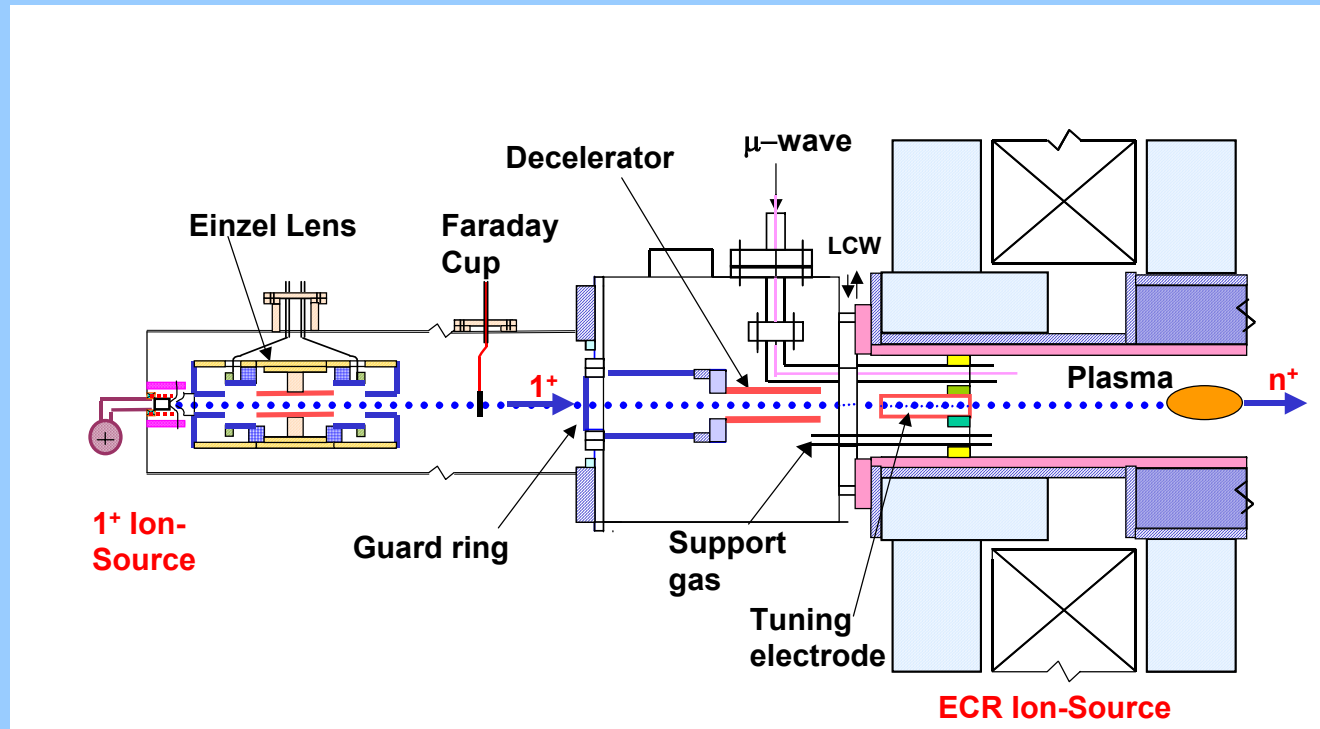
## R&D Initiative on ion source :

Integrated Target-ion source

1<sup>+</sup> RIB

ECR ion-source

n<sup>+</sup> RIB



Challenges

Decelerate 1<sup>+</sup> ions ~20 eV to ensure soft landing within ECRIS

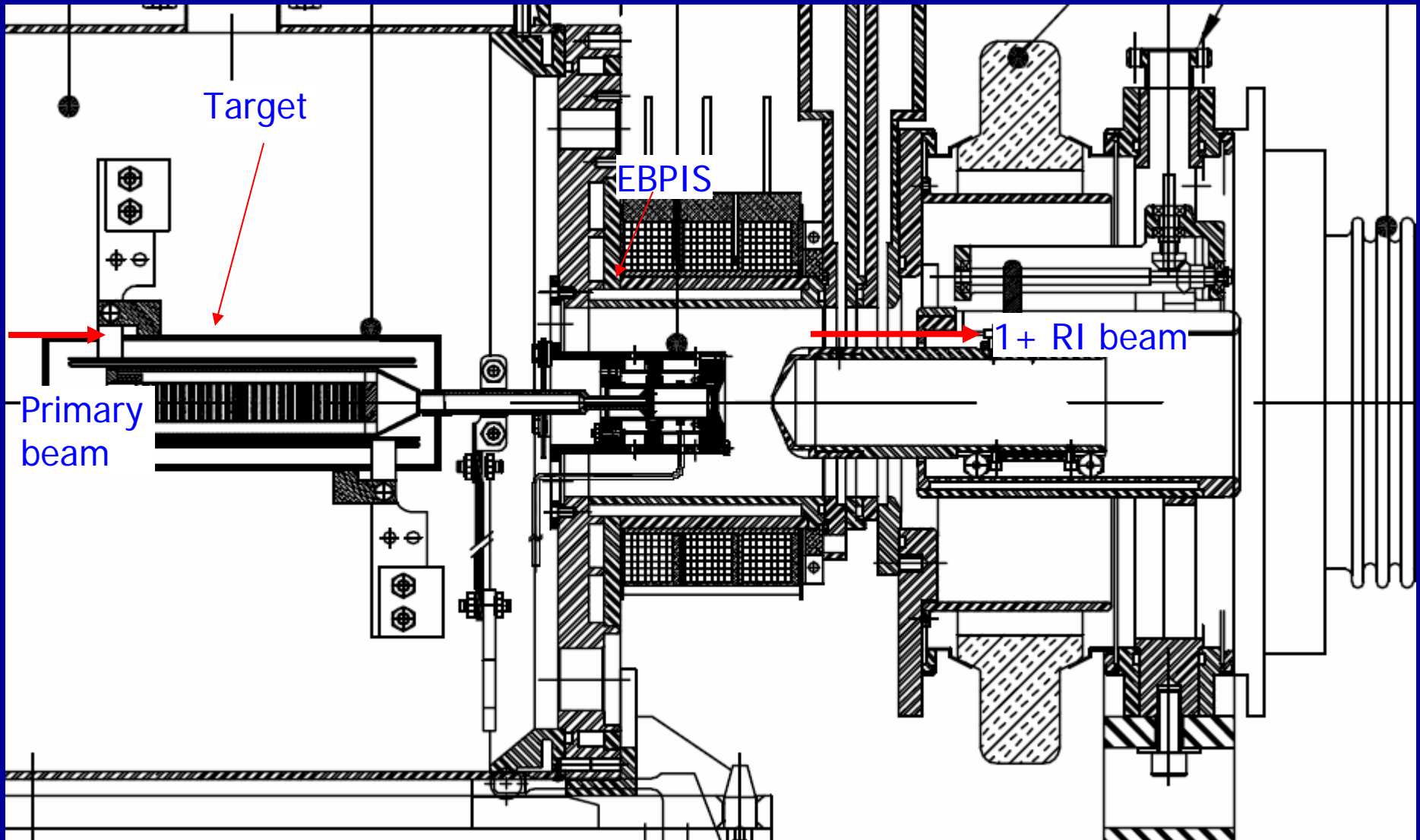
Optimised focussing to prevent beam loss

✓ ECRIS : at safe distance ~ 10<sup>-6</sup> mbar vacuum inside plasma chamber

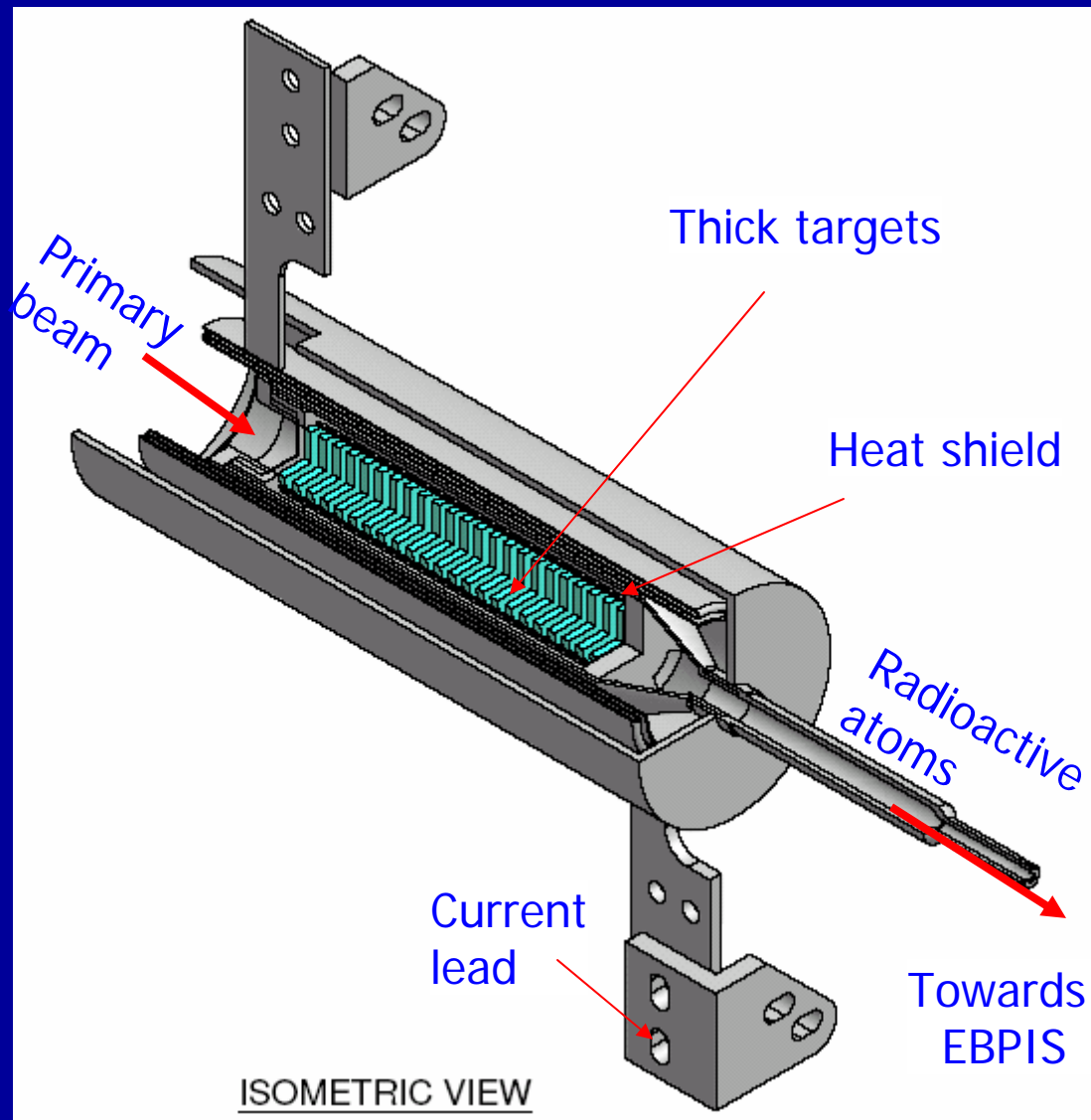
✓ ECR permanent magnets protected from high radiation near target

R&D Initiative on ion source :

## Integrated Thick-Target Electron Beam Plasma Ion-Source



## R&D Initiative on ion source :



Target holder

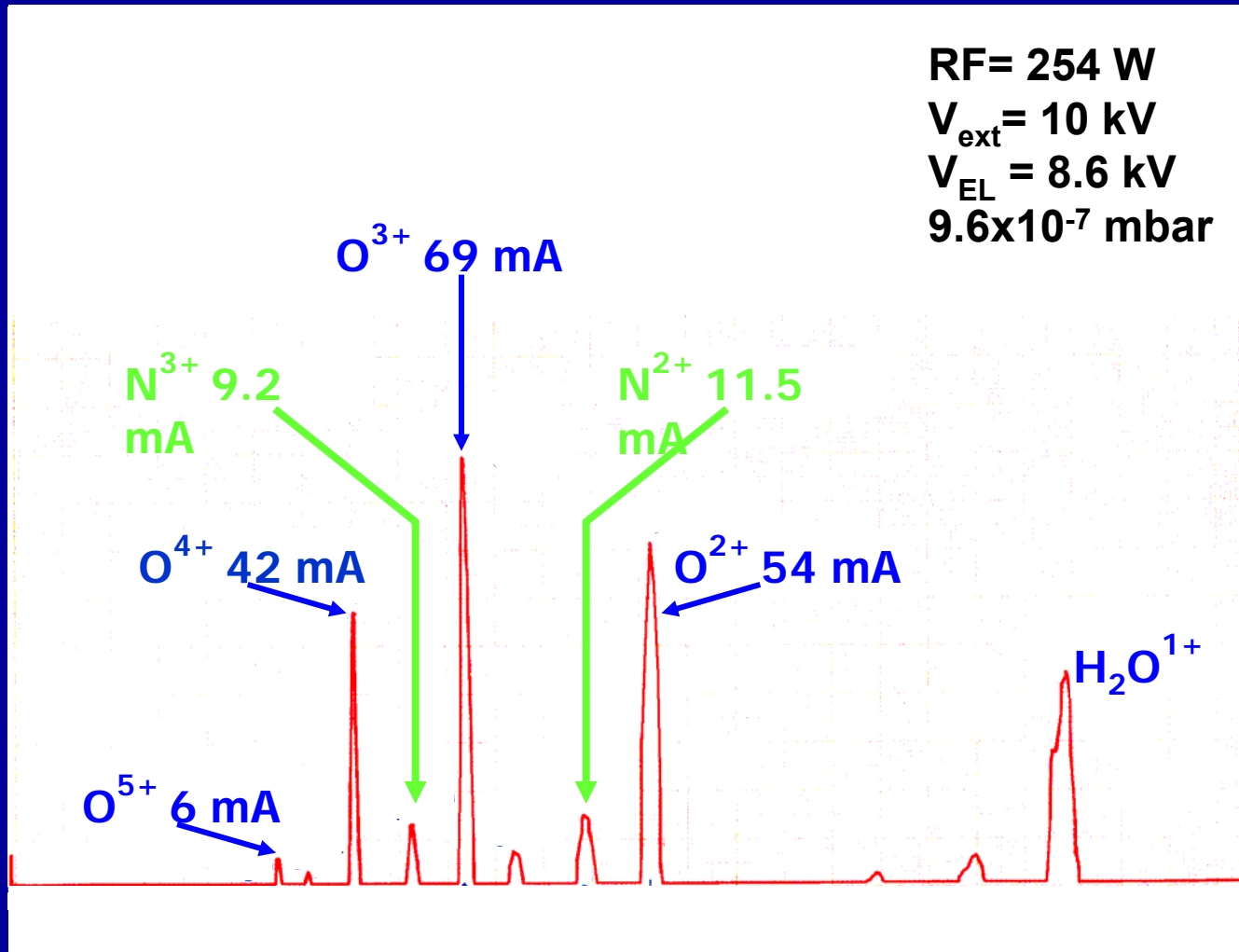


## R&D Initiative on ion source :



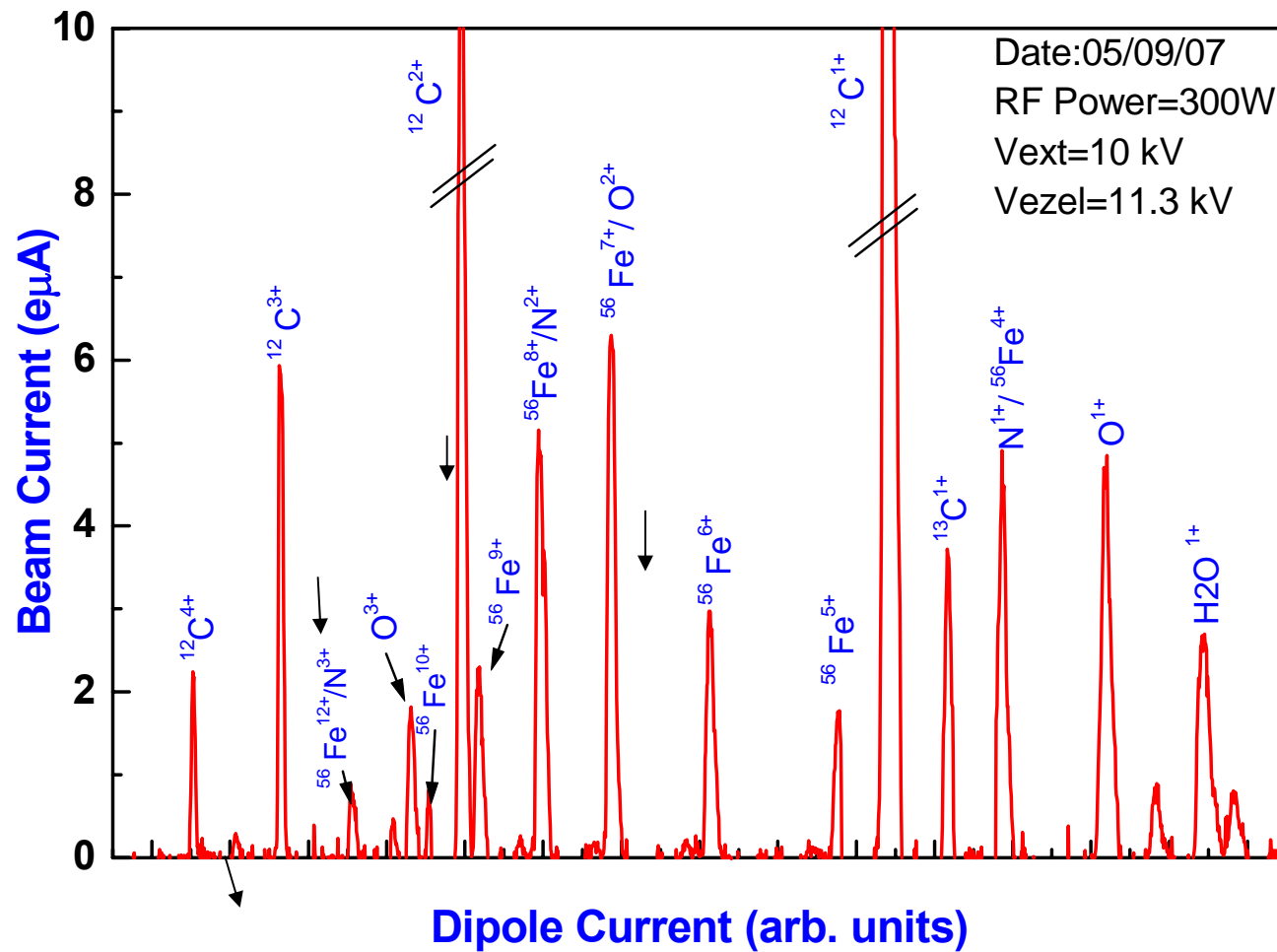
<b>Frequency</b>	<b>6.4 GHz</b>
<b>Klystron / Sol. Power</b>	<b>3 / 60 kW</b>
<b><math>B_{\text{ECR}}</math></b>	<b>0.23 T</b>
<b><math>B_z(\text{Inj}) / B_z(\text{Ext}) / B_r(r=R)</math></b>	<b>0.95 / 0.7 / 0.7 T</b>
<b>Mirror Ratio (Inj/Ext)</b>	<b>5.9 / 4.375</b>

# Typical Oxygen spectrum from ECR ion-source

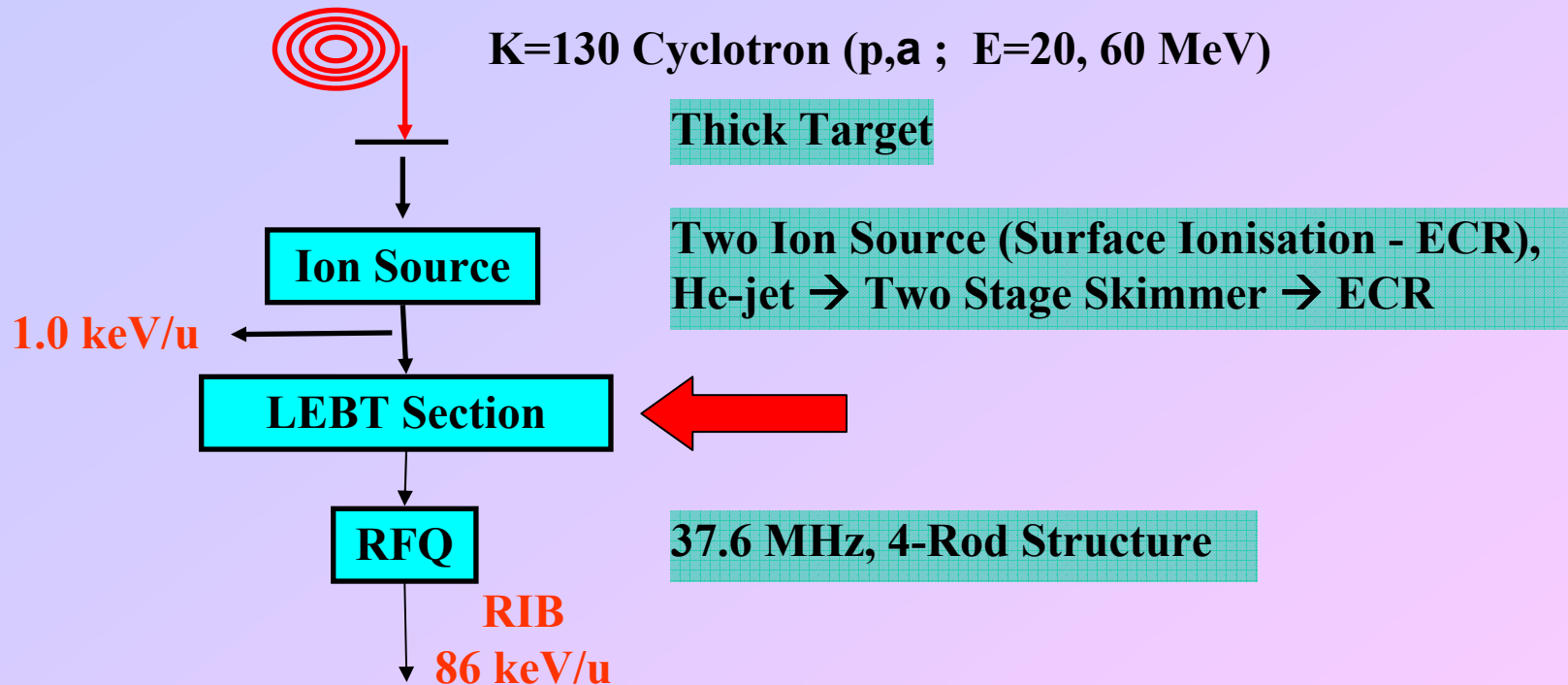




## First Iron spectrum



# Design & Development of LEBT section between on-line ECRIS & RFQ



## Design Aims :

- Initial separation of the RIB of interest having optimum  $q$  from the rest.
- Transverse matching of the selected RIB to the acceptance of the RFQ.

## Challenges :

- Maximum transmission & RP even with high emittance beam from IS (ECR).
- Minimum floor space & number of optical elements.

## Low Energy Beam Transport section

E  
C  
R  
I  
S

Analysing section

Focal Plane

Matching section

Solenoid length 0.25 m  
Solenoid Radius 0.65 m  
Max. Field 0.65 T

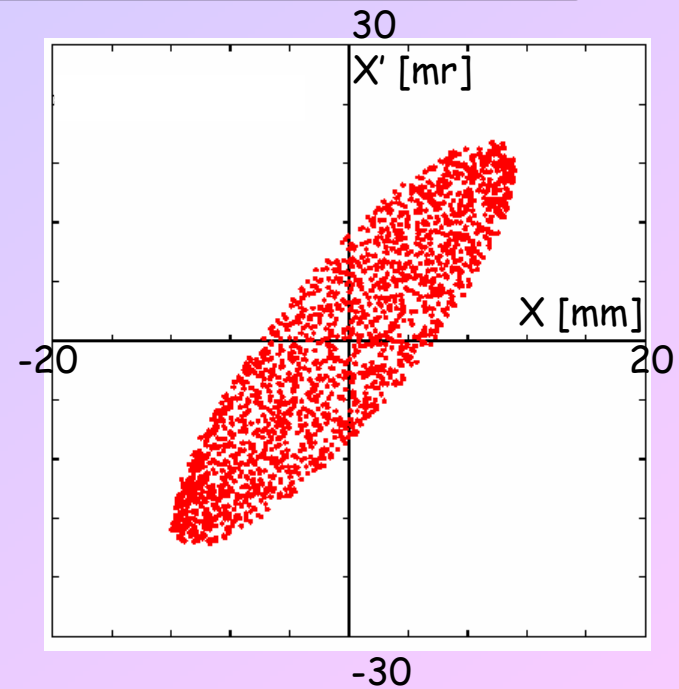
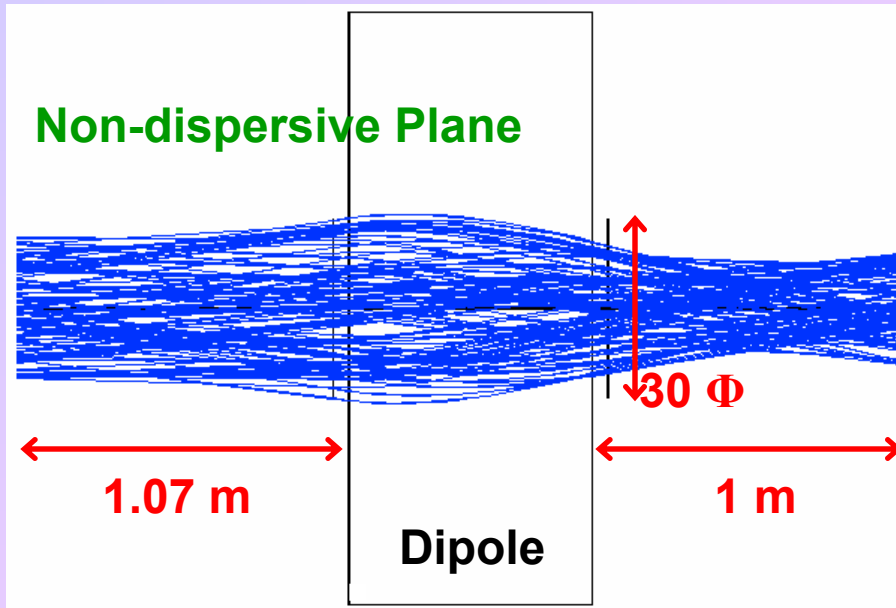
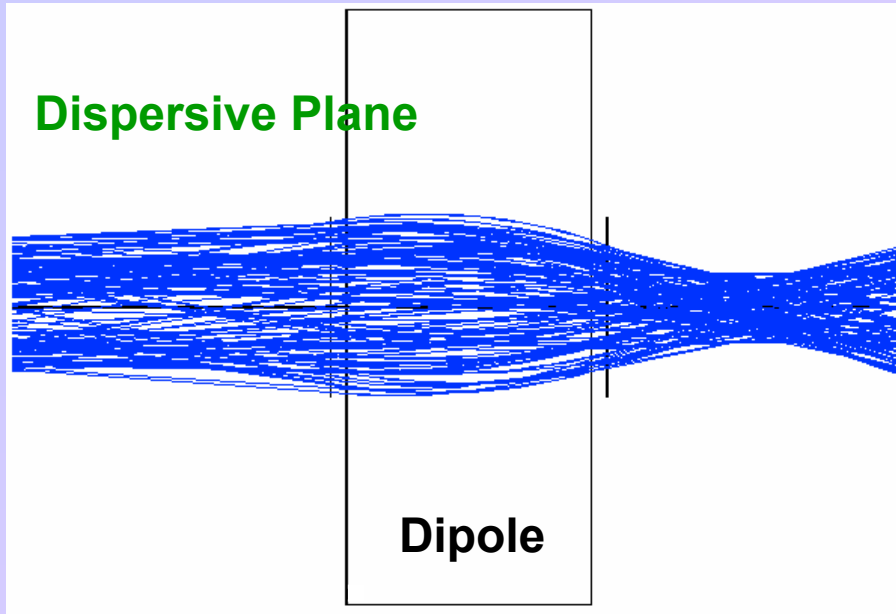
Virtual Source point

RFQ

Erect Ellipse  
 $e = 120 \pi \text{ mm mrad}$   
( $\pm 1.2 \text{ cm} / \pm 10 \text{ mrad}$ )

Bending Angle  $90^\circ$   
Bending Radius 0.5 m  
Max. Field 0.25 T  
Entry / Exit angle  $27.141^\circ / 27.107^\circ$

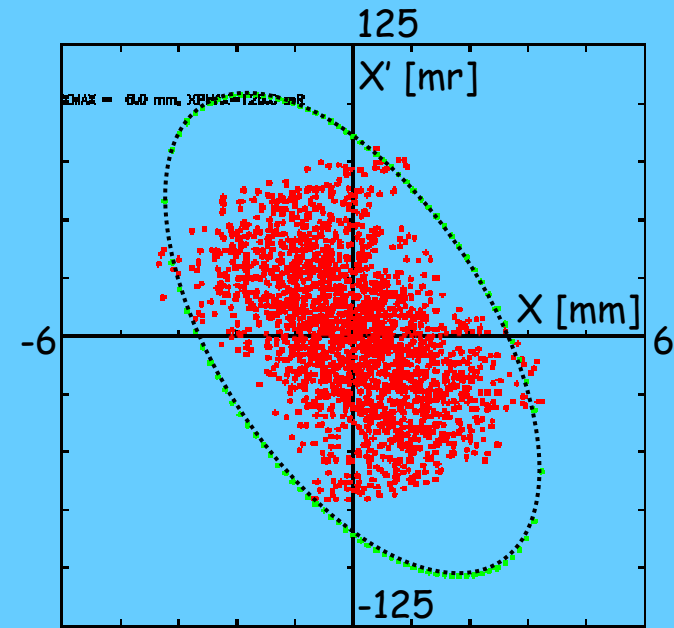
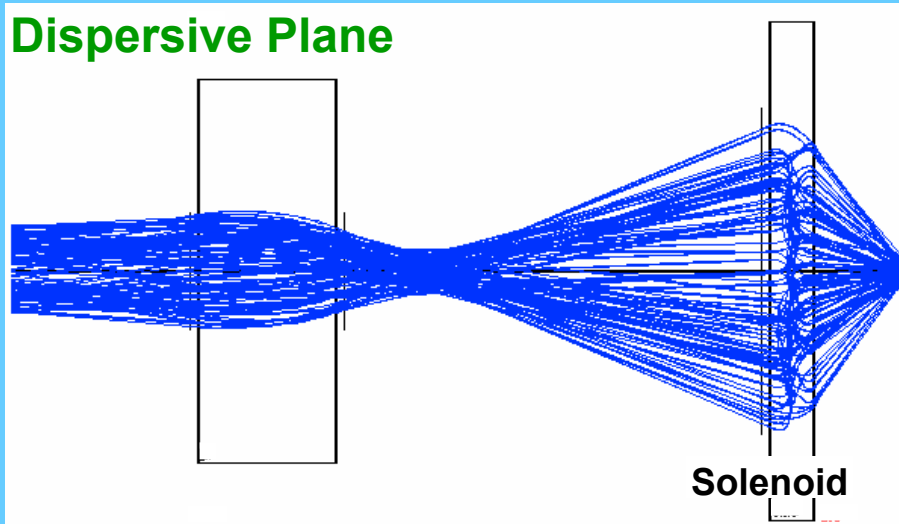
## Optics for the analysing section of the LEBT design



<b>Configuration</b>	<b>Drift-Dipole-Drift</b>
<b>Acceptance</b>	<b>120 <math>\pi</math>-mm-mr. (Full Transmission)</b>
<b>System Length</b>	<b>2.855 m</b>
<b>Dispersion</b>	<b>1.98 m (for 100% rigidity change)</b>
<b>Mass R. Power</b>	<b>43 (3<sup>rd</sup> order)</b>
<b>Magnification</b>	<b>-0.95</b>

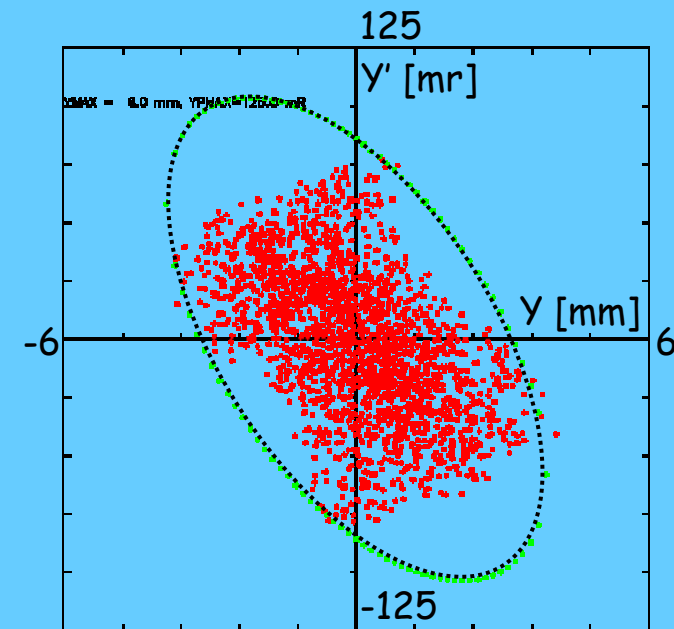
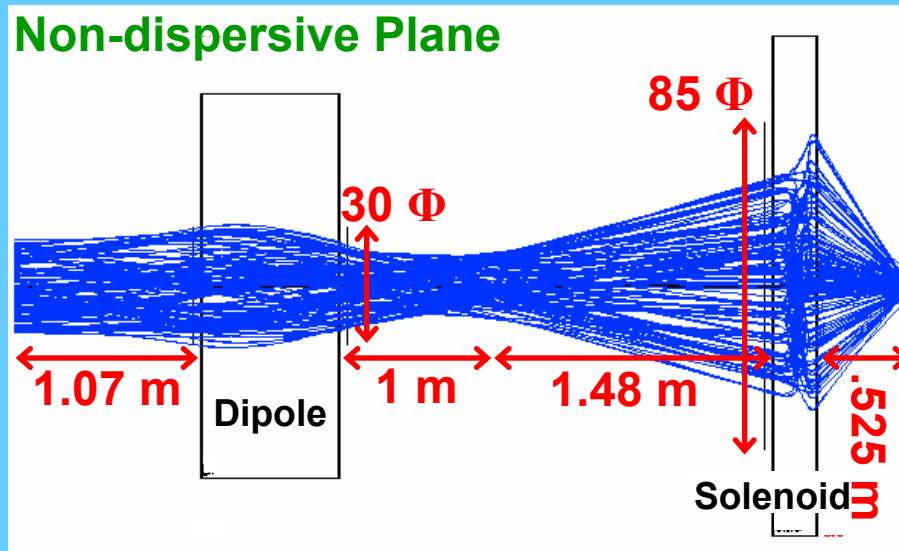
# Optics for the LEBT section

## Dispersive Plane



**Transmission 98.8 %**

## Non-dispersive Plane

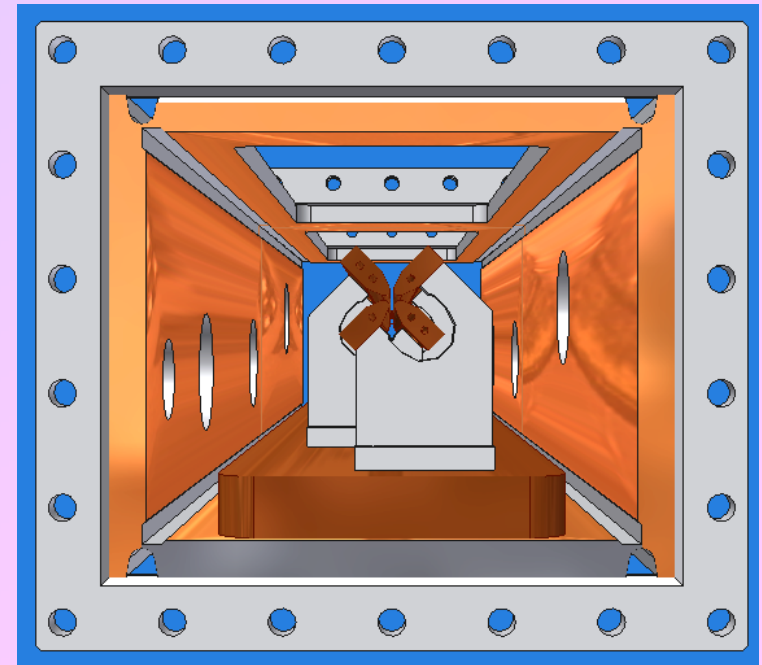
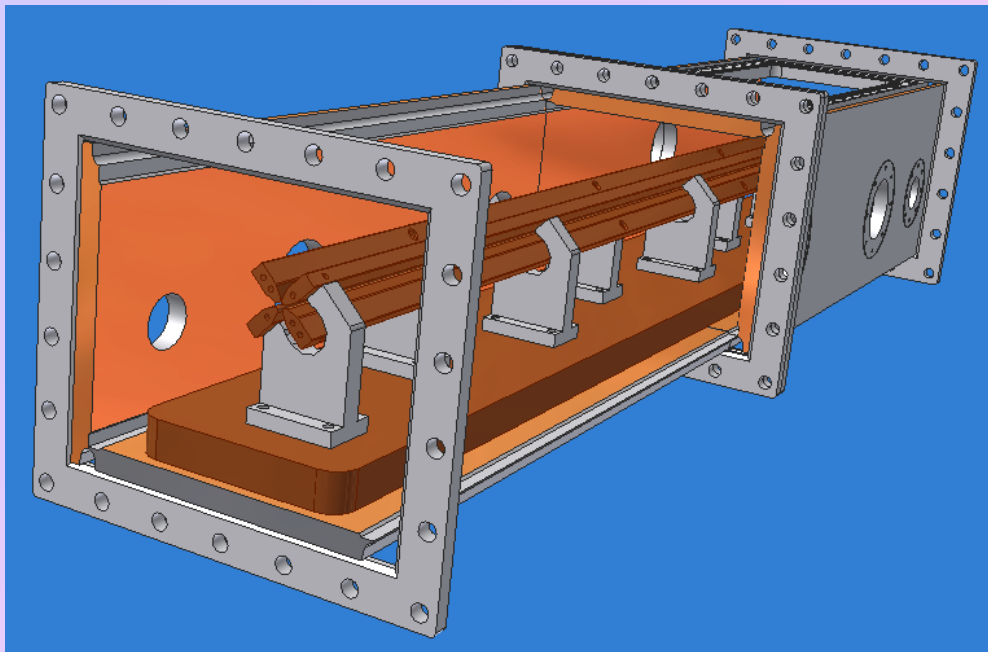


## Design of Radio Frequency Quadrupole (RFQ)

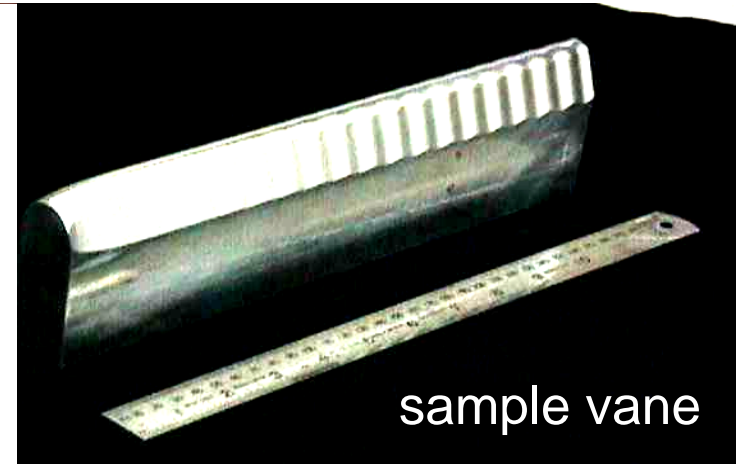
**Design Aim : Acceleration of RIBs ( $q/A \geq 1/16$ ) from 1 keV/u to about 90keV/u**

**RFQ : Single RF structure capable of bunching, focussing and accelerating low energy ( $\sim$  keV) ions**

**However  $\rightarrow$  Designing Machining and aligning**



## Vane mac



Vane profile from PARMTEQ



Co-ordinates of Surface



3D modelling of profile



Optimum tool path generation



Machining of vane



Check co-ordinates (CMM)  
&  
surface finish (Talysurf)

Requirement

$\Delta/r_0 \leq 0.5\%$  for good  
transmission :  $\Delta \sim 30\mu\text{m}$

Skin Depth  $d \sim 11\text{ mm}$   
 $R_t \leq 2.2\text{ mm}$

Sample vane

$\sim 20\text{ mm}$

$\sim 2\text{ mm}$



# RADIOFREQUENCY QUADRUPOLE (RFQ) : 1<sup>st</sup> in India

আনন্দবাজার পত্রিকা

১৫ আশ্বিন ১৪১২ শনিবার ১ অক্টোবর

**THE HINDU**

Online edition of India's National Newspaper  
Wednesday, Oct 05, 2005

সংক্ষেপে ...

পরমাণু বিজ্ঞানে নয়া সাফল্য ভারতের

স্টাফ রিপোর্টার ✧ কলকাতা

জাপানের পরে এ বার ভারতেও রেডিও ফ্রিকোয়েন্সি কোয়ার্ট্র  
পল চালু হল। এশিয়ার মধ্যে ভারতই হল দ্বিতীয় দেশ,  
যেখানে এই 'অ্যাক্সিলারেটর' বা ত্বারক চালু করা হয়েছে।  
শুক্রবার ডেরিয়েবল এনার্জি সাইক্লোট্রন সেন্টারের অধিকর্তা  
বিকাশ সিংহ এক লিখিত বিবৃতিতে এ কথা জানান। এটি একটি  
জটিল এবং অত্যধিক ব্যয়বহুল ত্বারক। এর মাধ্যমে পরমাণু কণাকে  
প্রচণ্ড গতিশীল করে তোলা যাবে। ১৯৮০ সালেই প্রথম মার্কিন  
যু.স.রা এই ত্বারক চালু করে। তার পর থেকে খুব বেশি দেশ  
এই ত্বারক চালু করতে পারেনি।

## National

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National

**India joins select club in particle technology**

Special Correspondent

KOLKATA: India's first heavy ion Radio Frequency Quadrupole [RFQ] accelerator has been commissioned at the Department of Atomic Energy's Variable Energy Cyclotron Centre [VECC] here.

Scientists from across the world have acknowledged the achievement as a hall-mark development in particle accelerator technology in the country, VECC officials told *The Hindu* on Tuesday.

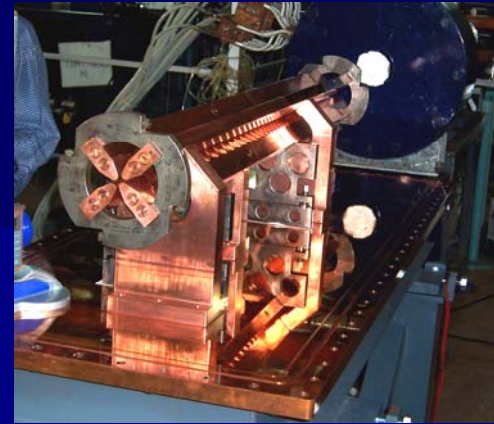
Japan is the only other Asian country to have successfully ... for which was tried out on a ... st time in the United States of

[z] cavity of very pure copper ... ed vanes which takes care of ... using of ion beams", according

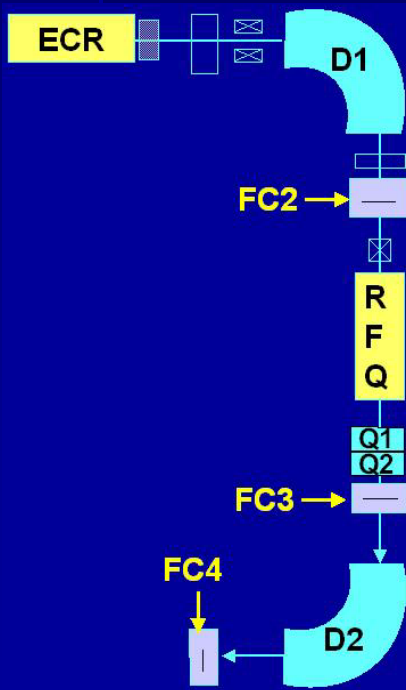


RFQ during installation





**Experimental Results: Test Beam: Ar 4+**



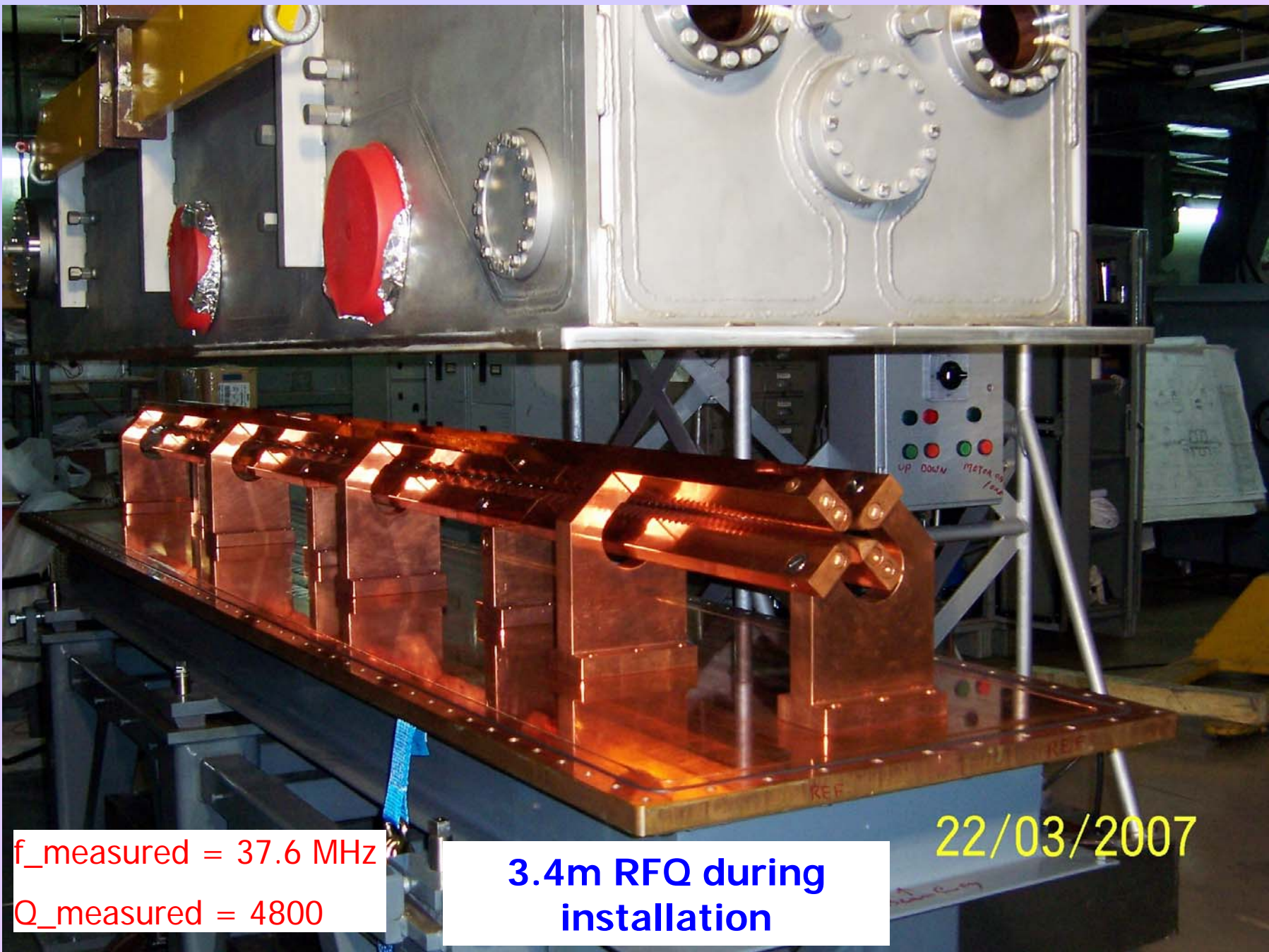
Calculated magnetic strength (kG) (29.06 keV/u)	Experimental magnetic strength (kG)	Transmission Efficiency* % *with electron suppression
Q1: -1.528 Q2: 1.065	Q1: -1.53 Q2: 1.057	~81% <b>FC3/FC2</b>
Q1 : -1.07 Q2: 0.7 D2 : 2.058	Q1: -1.058 Q2: 0.82 D2: 2.058	~80% <b>FC4/FC2</b>

## Beams available from the RIB facility

- ◆ Oxygen : up to 120 keV (after ECR); 464 keV (after RFQ)
- ◆ Nitrogen : up to 100 keV (after ECR); 406 keV (after RFQ)
- ◆ Argon : up to 160 keV (after ECR); 1.16 MeV (after RFQ)
- ◆ Iron : up to 220 keV (after ECR); 1.6 MeV (after RFQ)
- ◆ Also H, Helium, O<sub>2</sub>, Carbon, ...

Typical measured currents: O<sup>3+</sup> ~ 70 μA; O<sup>4+</sup> 40 μA; O<sup>5+</sup> ~ 6 μA;  
Ar<sup>4+</sup> ~ 4 μA; He<sup>1+</sup> ~ 100 μA; Fe<sup>6+</sup> ~ 7 μA; Fe<sup>10+</sup> ~ 1 μA

*optimization of ECR continuing*

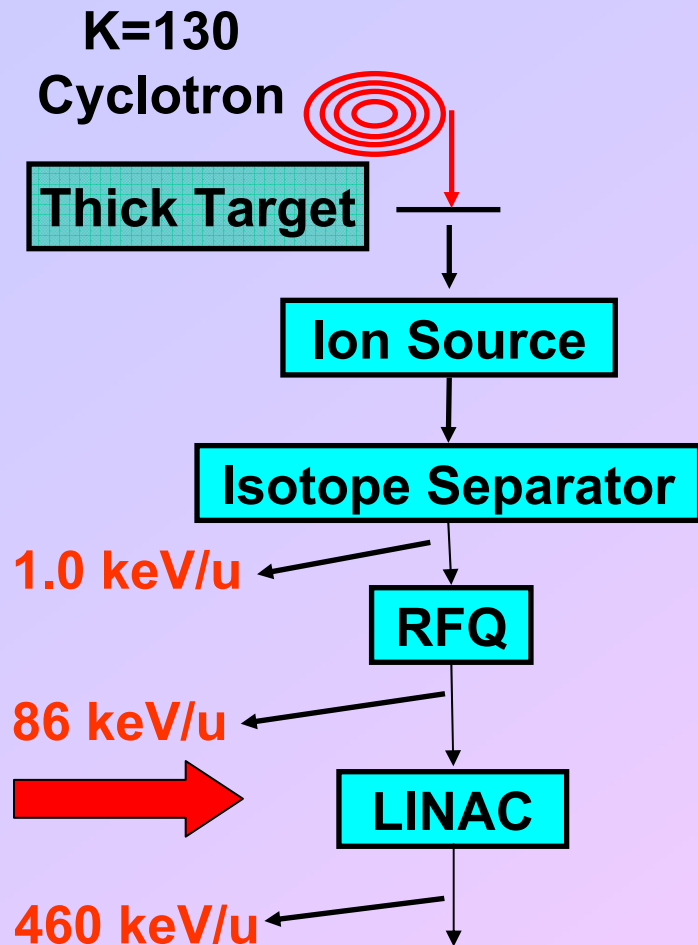


$f_{\text{measured}} = 37.6 \text{ MHz}$   
 $Q_{\text{measured}} = 4800$

**3.4m RFQ during  
installation**

22/03/2007

## Design of LINAC cavities



### Design Parameters

- $q/A \geq 1/14$
- $e_n \geq 0.5$  p-mm-mr
- $T_{in} = 86$  keV/u

### Design goals

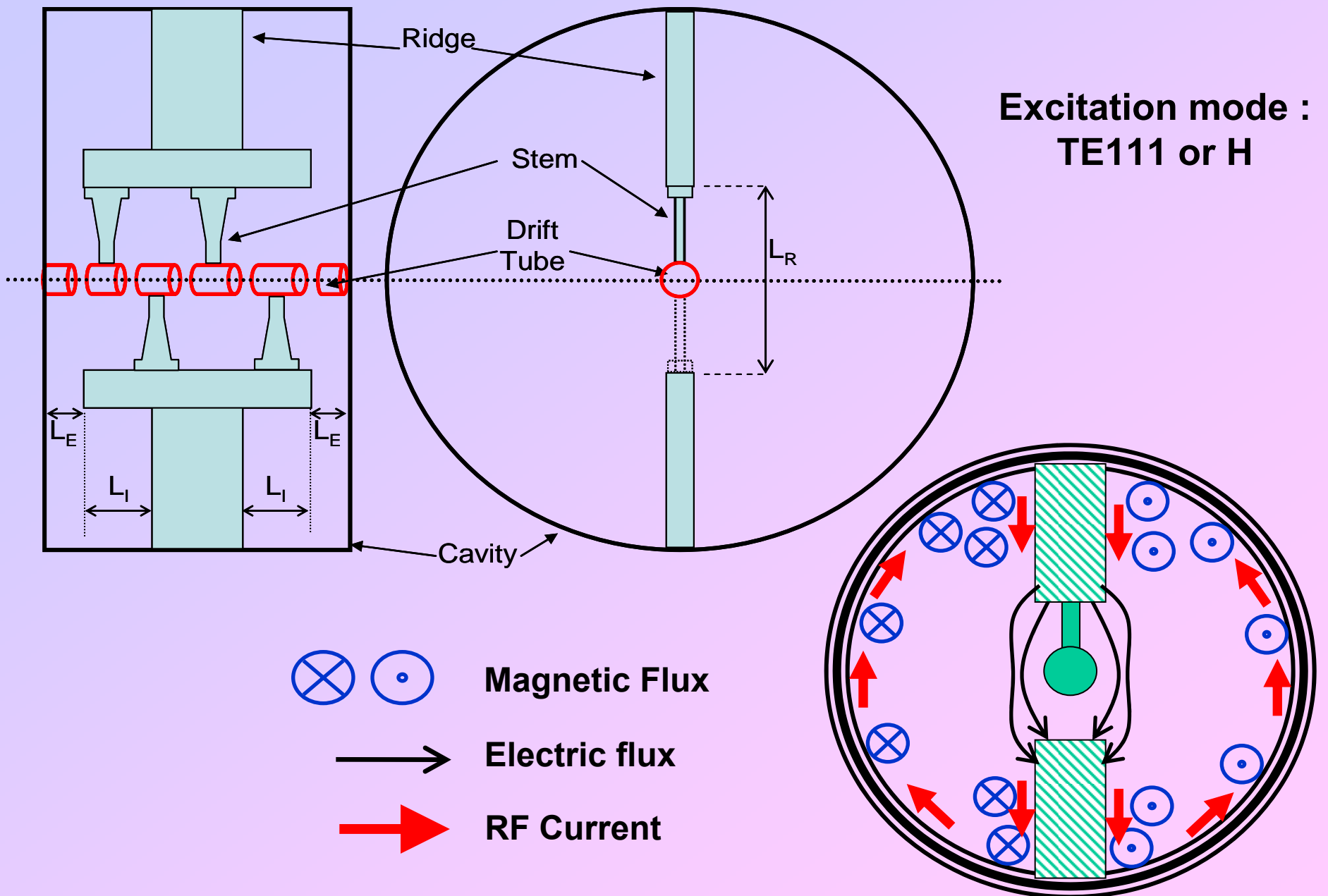
#### Beam Dynamics :

- Maximising Transmission
- Energy tunability
- Good beam quality (DE, Dt)

#### RF Analysis :

- Getting the desired frequency
- Optimisation for best shunt impedance
- Surface current density at the junction of two components should not be high

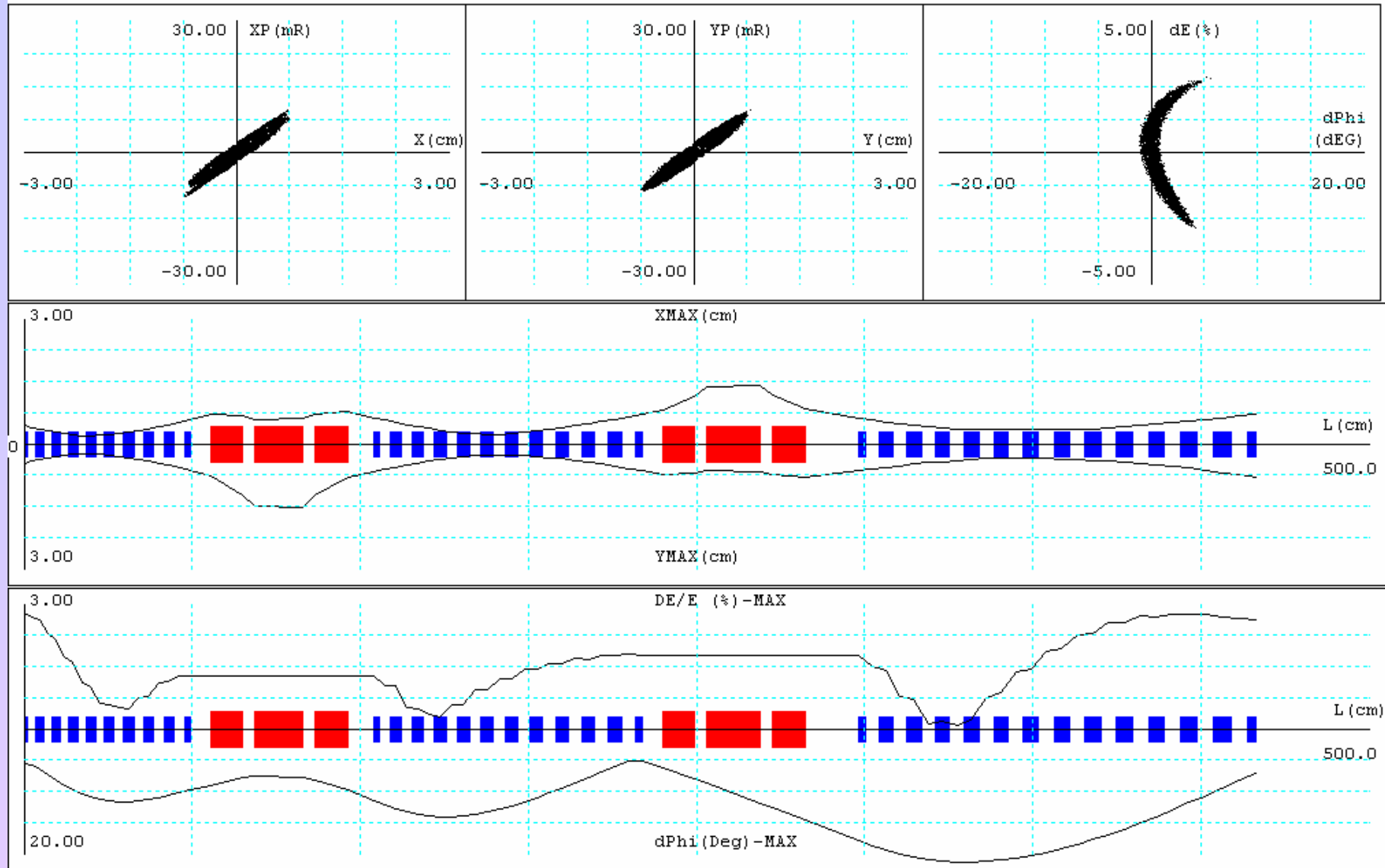
# IH - LINAC :





# VECLIN Snap shot

$\epsilon_{x/y} = 0.5 \pi$  mm mrad &  $\epsilon_z = 10 \pi$ -deg-% A=16



Inter-tank space 75 cm

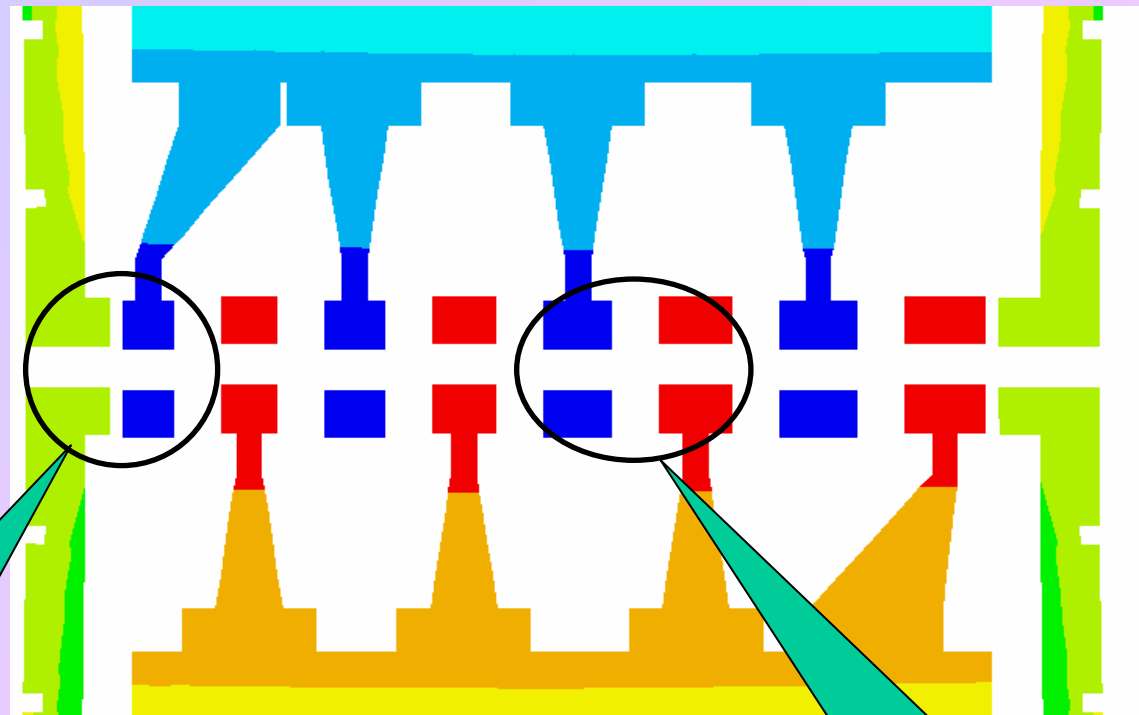
Quads  $\rightarrow$  L=21+14+21 cm / Aperture rad 2.85 cm / Gradient 30 T/m

# IH - LINAC : The design details

## 6. Engineering Analysis ANSYS

Structural deflection under various loads :

(a) Atmospheric pressure (b) Self weight (c) Thermal flux due to RF



**Axial  
deflection of  
flange**

**Radial  
deflection of  
cavity**

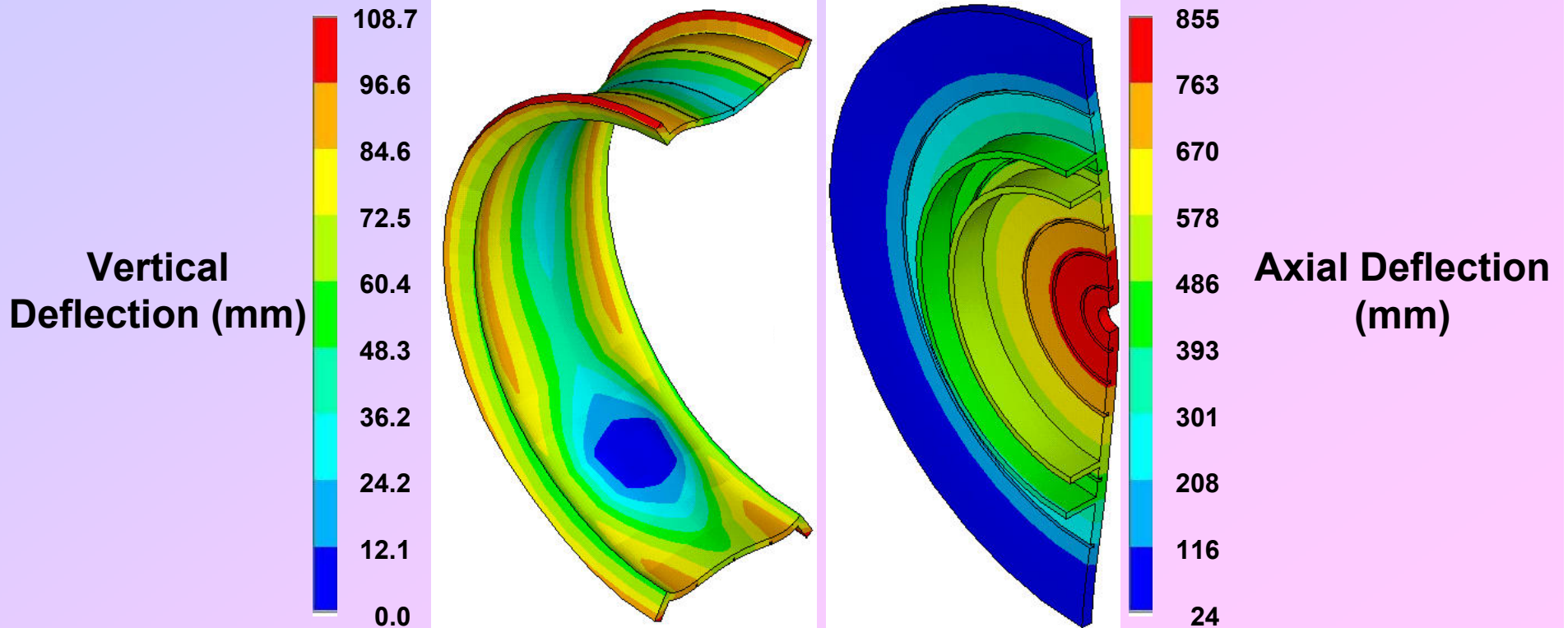
The effects have been tried to minimise using :

(a) Proper thickness of the cavity materials :

Mild Steel – 25 mm (cavity) & 35 mm (End covers)

(b) Circular stiffeners on end covers : Two Nos.

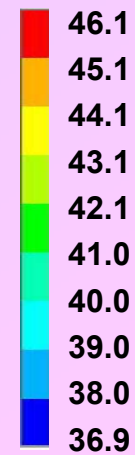
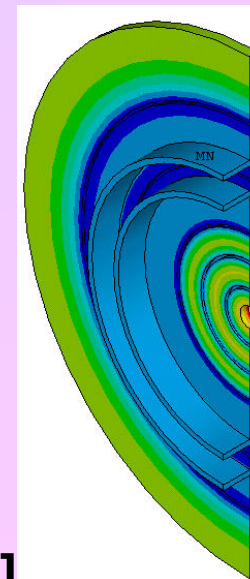
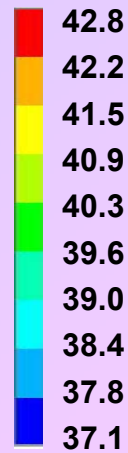
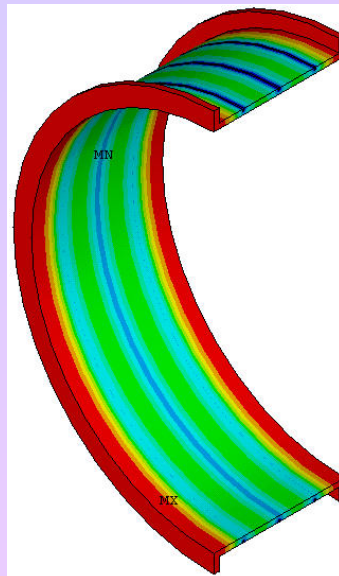
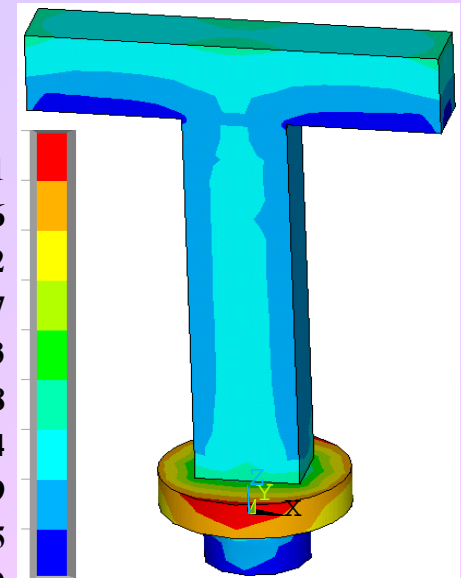
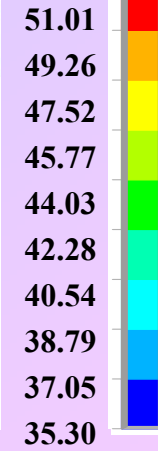
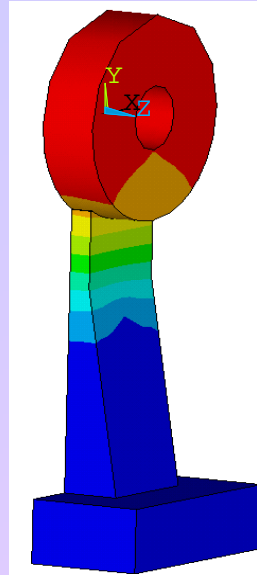
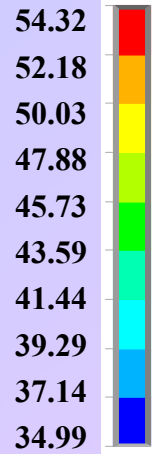
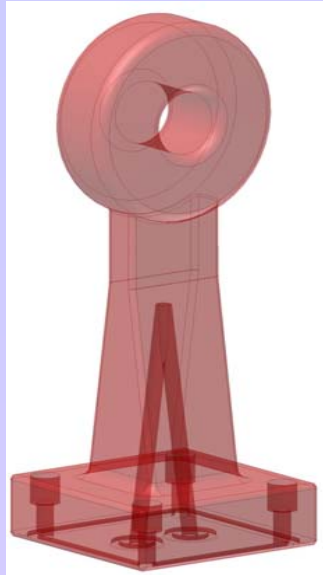
(c) Cooling





# IH - LINAC : The design details

# Temperature Distribution

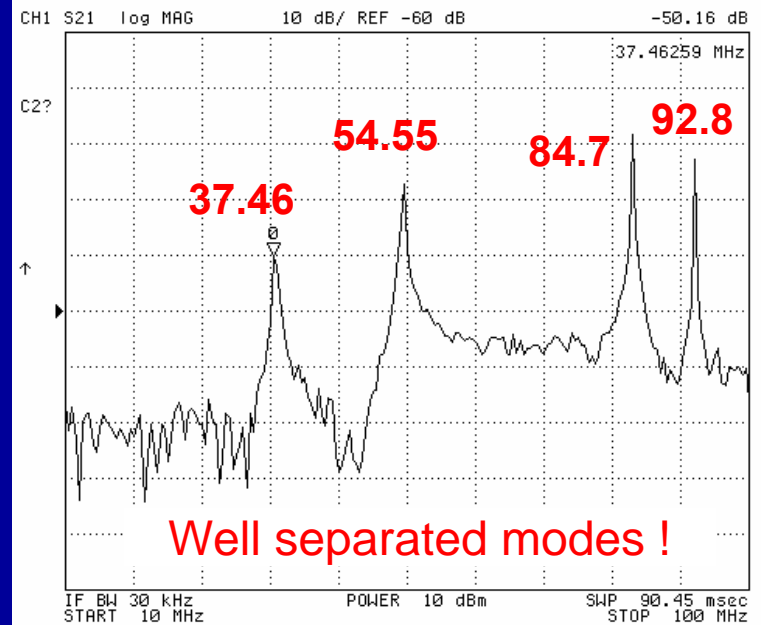
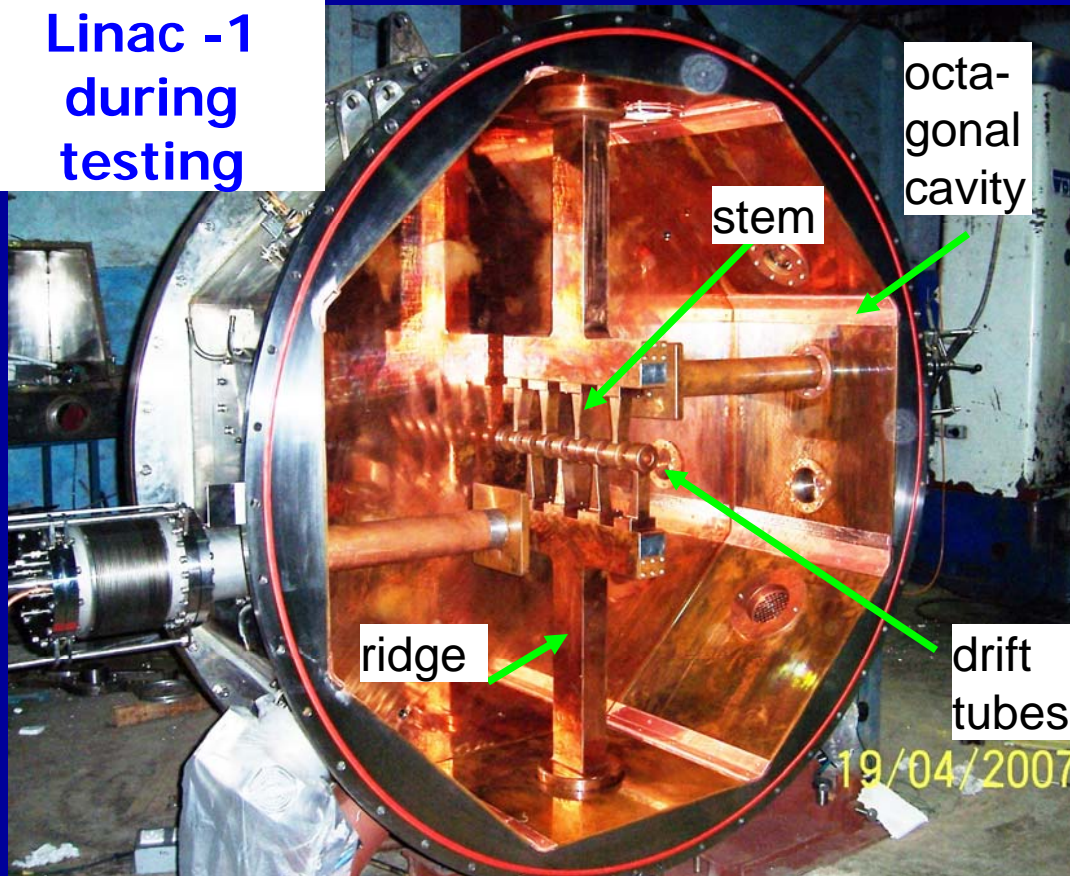


Temperature [°C]

## LINAC-1 : Important Parameters

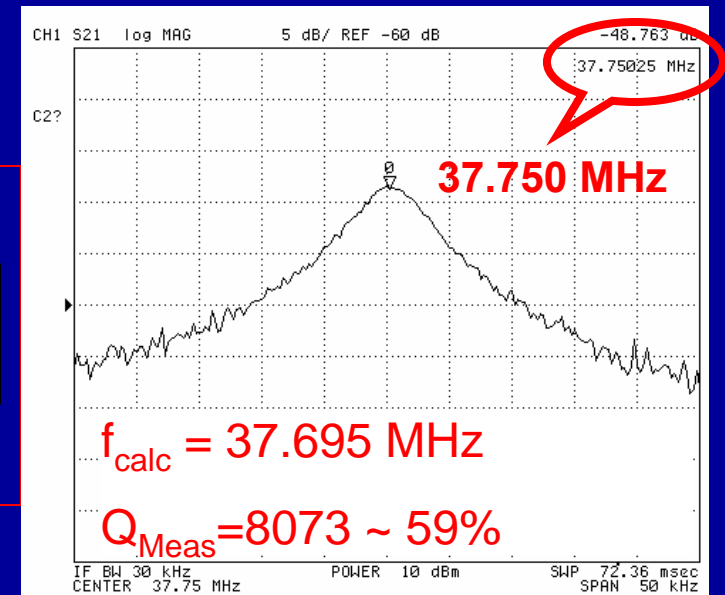
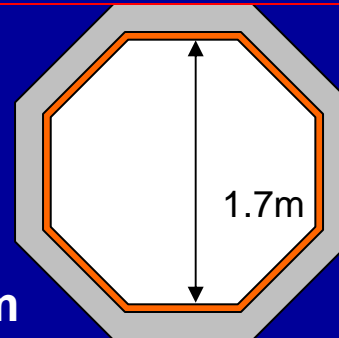
Frequency	MHz	37.6
q/A		$\geq 1/14$
$T_{in} \rightarrow T_{out}$	keV/u	98.785 $\rightarrow$ 183.56
$\beta_{in} \rightarrow \beta_{out}$	%	1.456 $\rightarrow$ 1.985
Accelerating gaps	#	9
Drift tube I/D & O/D	mm	25 & 69.5
Drift tube gap	mm	29.2
Cell length	mm	58.4 $\rightarrow$ 78.8
Peak D.T. voltage	kV	101.24 kV
Max. Field		1.4 x $E_{Kilpatrick}$
Transit Factor		0.789 $\rightarrow$ 0.843
Sync. Phase	Degree	-24
Cavity Length	M	0.6182
Accln. Gradient	MV/m per q	2.102
Shunt Impedance	M $\Omega$ /m	342
Q-Value		13765
Power	kW	$\sim 15$

# Linac -1 during testing



## CAVITY :

- 25 mm thk. SS304L
- clad with 5 mm thk. Cu
- Diameter ~ 1.7m ; Length ~ 0.6 m







VACUUM TECHNOLOGIES PVT LTD  
100, 101, 102, 103, 104, 105, 106, 107, 108, 109, 110, 111, 112, 113, 114, 115, 116, 117, 118, 119, 120, 121, 122, 123, 124, 125, 126, 127, 128, 129, 130, 131, 132, 133, 134, 135, 136, 137, 138, 139, 140, 141, 142, 143, 144, 145, 146, 147, 148, 149, 150, 151, 152, 153, 154, 155, 156, 157, 158, 159, 160, 161, 162, 163, 164, 165, 166, 167, 168, 169, 170, 171, 172, 173, 174, 175, 176, 177, 178, 179, 180, 181, 182, 183, 184, 185, 186, 187, 188, 189, 190, 191, 192, 193, 194, 195, 196, 197, 198, 199, 200, 201, 202, 203, 204, 205, 206, 207, 208, 209, 210, 211, 212, 213, 214, 215, 216, 217, 218, 219, 220, 221, 222, 223, 224, 225, 226, 227, 228, 229, 230, 231, 232, 233, 234, 235, 236, 237, 238, 239, 240, 241, 242, 243, 244, 245, 246, 247, 248, 249, 250, 251, 252, 253, 254, 255, 256, 257, 258, 259, 260, 261, 262, 263, 264, 265, 266, 267, 268, 269, 270, 271, 272, 273, 274, 275, 276, 277, 278, 279, 280, 281, 282, 283, 284, 285, 286, 287, 288, 289, 290, 291, 292, 293, 294, 295, 296, 297, 298, 299, 300, 301, 302, 303, 304, 305, 306, 307, 308, 309, 310, 311, 312, 313, 314, 315, 316, 317, 318, 319, 320, 321, 322, 323, 324, 325, 326, 327, 328, 329, 330, 331, 332, 333, 334, 335, 336, 337, 338, 339, 340, 341, 342, 343, 344, 345, 346, 347, 348, 349, 350, 351, 352, 353, 354, 355, 356, 357, 358, 359, 360, 361, 362, 363, 364, 365, 366, 367, 368, 369, 370, 371, 372, 373, 374, 375, 376, 377, 378, 379, 380, 381, 382, 383, 384, 385, 386, 387, 388, 389, 390, 391, 392, 393, 394, 395, 396, 397, 398, 399, 400, 401, 402, 403, 404, 405, 406, 407, 408, 409, 410, 411, 412, 413, 414, 415, 416, 417, 418, 419, 420, 421, 422, 423, 424, 425, 426, 427, 428, 429, 430, 431, 432, 433, 434, 435, 436, 437, 438, 439, 440, 441, 442, 443, 444, 445, 446, 447, 448, 449, 450, 451, 452, 453, 454, 455, 456, 457, 458, 459, 460, 461, 462, 463, 464, 465, 466, 467, 468, 469, 470, 471, 472, 473, 474, 475, 476, 477, 478, 479, 480, 481, 482, 483, 484, 485, 486, 487, 488, 489, 490, 491, 492, 493, 494, 495, 496, 497, 498, 499, 500, 501, 502, 503, 504, 505, 506, 507, 508, 509, 510, 511, 512, 513, 514, 515, 516, 517, 518, 519, 520, 521, 522, 523, 524, 525, 526, 527, 528, 529, 530, 531, 532, 533, 534, 535, 536, 537, 538, 539, 540, 541, 542, 543, 544, 545, 546, 547, 548, 549, 550, 551, 552, 553, 554, 555, 556, 557, 558, 559, 560, 561, 562, 563, 564, 565, 566, 567, 568, 569, 570, 571, 572, 573, 574, 575, 576, 577, 578, 579, 580, 581, 582, 583, 584, 585, 586, 587, 588, 589, 590, 591, 592, 593, 594, 595, 596, 597, 598, 599, 600, 601, 602, 603, 604, 605, 606, 607, 608, 609, 610, 611, 612, 613, 614, 615, 616, 617, 618, 619, 620, 621, 622, 623, 624, 625, 626, 627, 628, 629, 630, 631, 632, 633, 634, 635, 636, 637, 638, 639, 640, 641, 642, 643, 644, 645, 646, 647, 648, 649, 650, 651, 652, 653, 654, 655, 656, 657, 658, 659, 660, 661, 662, 663, 664, 665, 666, 667, 668, 669, 670, 671, 672, 673, 674, 675, 676, 677, 678, 679, 680, 681, 682, 683, 684, 685, 686, 687, 688, 689, 690, 691, 692, 693, 694, 695, 696, 697, 698, 699, 700, 701, 702, 703, 704, 705, 706, 707, 708, 709, 710, 711, 712, 713, 714, 715, 716, 717, 718, 719, 720, 721, 722, 723, 724, 725, 726, 727, 728, 729, 730, 731, 732, 733, 734, 735, 736, 737, 738, 739, 740, 741, 742, 743, 744, 745, 746, 747, 748, 749, 750, 751, 752, 753, 754, 755, 756, 757, 758, 759, 760, 761, 762, 763, 764, 765, 766, 767, 768, 769, 770, 771, 772, 773, 774, 775, 776, 777, 778, 779, 780, 781, 782, 783, 784, 785, 786, 787, 788, 789, 790, 791, 792, 793, 794, 795, 796, 797, 798, 799, 800, 801, 802, 803, 804, 805, 806, 807, 808, 809, 810, 811, 812, 813, 814, 815, 816, 817, 818, 819, 820, 821, 822, 823, 824, 825, 826, 827, 828, 829, 830, 831, 832, 833, 834, 835, 836, 837, 838, 839, 840, 841, 842, 843, 844, 845, 846, 847, 848, 849, 850, 851, 852, 853, 854, 855, 856, 857, 858, 859, 860, 861, 862, 863, 864, 865, 866, 867, 868, 869, 870, 871, 872, 873, 874, 875, 876, 877, 878, 879, 880, 881, 882, 883, 884, 885, 886, 887, 888, 889, 890, 891, 892, 893, 894, 895, 896, 897, 898, 899, 900, 901, 902, 903, 904, 905, 906, 907, 908, 909, 910, 911, 912, 913, 914, 915, 916, 917, 918, 919, 920, 921, 922, 923, 924, 925, 926, 927, 928, 929, 930, 931, 932, 933, 934, 935, 936, 937, 938, 939, 940, 941, 942, 943, 944, 945, 946, 947, 948, 949, 950, 951, 952, 953, 954, 955, 956, 957, 958, 959, 960, 961, 962, 963, 964, 965, 966, 967, 968, 969, 970, 971, 972, 973, 974, 975, 976, 977, 978, 979, 980, 981, 982, 983, 984, 985, 986, 987, 988, 989, 990, 991, 992, 993, 994, 995, 996, 997, 998, 999, 1000

03/12/2007

## LINAC-2 : Important Parameters

Frequency	MHz	37.6
q/A		$\geq 1/14$
$T_{in} \rightarrow T_{out}$	keV/u	183.56 $\rightarrow$ 286.8
$\beta_{in} \rightarrow \beta_{out}$	%	1.985 $\rightarrow$ 2.481
Accelerating gaps	#	10
Drift tube I/D & O/D	mm	25 & 60
Drift tube gap	mm	39.8
Cell length	mm	79.6 $\rightarrow$ 98.4
Peak D.T. voltage	kV	107.5 kV
Max. Field		1.3 x $E_{Kilpatrick}$
Transit Factor		0.8045 $\rightarrow$ 0.8506
Sync. Phase	Degree	-25
Cavity Length (Inner)	M	0.871
Accln. Gradient	MV/m per q	1.79
Shunt Impedance	M $\Omega$ /m	432
Q-Value (Calc.)		18856
Power (Calc.)	kW	$\sim 10$





## First few beams →

RIB	T1/2	Production	Target
$^{11}\text{C}$	20 min	$^{11}\text{B}(p,n)$	BN
$^{13}\text{N}$	10 min	$^{13}\text{C}(p,n)$	Graphite
$^{17}\text{F}$	1 min	$^{14}\text{N}(\alpha,n)$	BN
$^{18}\text{F}$	110 min	$^{16}\text{O}(\alpha,n)$	$\text{HfO}_2, \text{Al}_2\text{O}_3$
$^{19}\text{Ne}$	17 sec	$^{19}\text{F}(p,n)$	LiF
$^{35}\text{Ar}$	1.7 sec	$^{35}\text{Cl}(p,n)$	$\text{CaCl}_2$
$^{38}\text{K}$	7.6 min	$^{35}\text{Cl}(\alpha,n)$	$\text{CaCl}_2$
$^{90}\text{Kr}$	32 sec	U/Th( $\alpha,f$ )	UC/ThO
$^{93}\text{Rb}$	6 sec	-do-	-do-

$$\begin{aligned}
 I_{\text{RIB}} &= I_{\text{Pri}} \cdot N_t \cdot \sigma \cdot \eta_{\text{release}} \cdot \eta_{\text{IS}} \cdot \eta_{\text{separation}} \cdot \eta_{\text{transport}} \\
 &= \frac{10^{-6}}{(1.6 \cdot 10^{-19})} \cdot (6 \cdot 10^{23}) \cdot (50 \cdot 10^{-3} \cdot 10^{-24}) \cdot 0.1 \cdot 0.05 \cdot 0.6 \cdot 0.5 \approx 10^9 \text{ pps} \\
 &\quad \underbrace{\hspace{10em}}_{0.15\%}
 \end{aligned}$$

$1\text{mA}$ 
 $1\text{g/cm}^2$ 
 $50 \text{ mb}$

Expected average yield at experimental station  $\sim 10^6\text{-}10^8$

**Future Plans**



## LINAC : What Next ???

**LINAC-1 : 98.785 → 183.56 keV/u**

- Installation at VECC have started : To be completed by 15<sup>th</sup> Oct, 2007
- Installation in beam line : April 2008

**LINAC-2 : 183.6 → 286.8 keV/u**

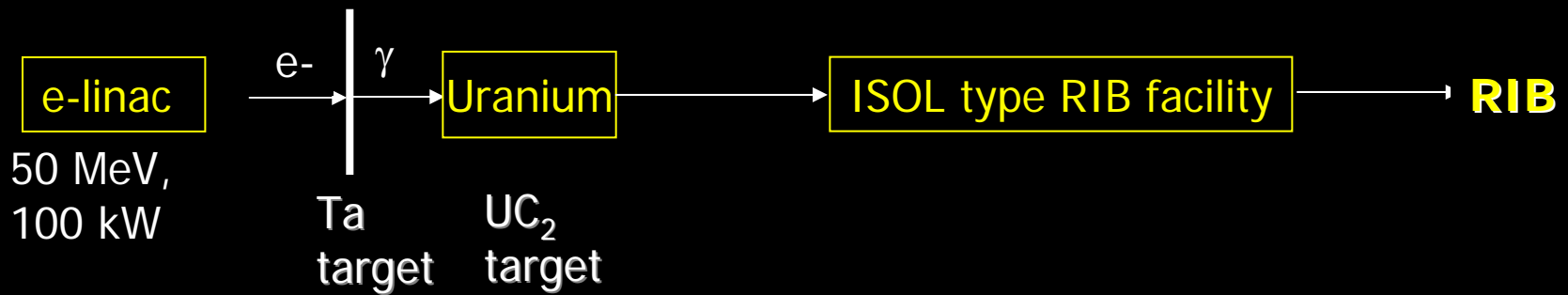
- PO placed
- Delivery schedule : Mar 2008
- Installation in beam line : Sep 2008

**LINAC-3 : 286.8 → 450 keV/u**

- Advanced stage of design
- Likely completion date : Dec 2008

**LINAC-4 to LINAC-8 : (To be completed by 2012)**

- Frequency 75.2 MHz
- 450 keV/u to 1.3 MeV/u
- Physics design stage



$\gamma$ - induced fission (photo-fission); GDR peak at  $\sim 13.5$  MeV  
 average photo-fission cross-section = 160 mb

Expected yield of some very neutron-rich exotic nuclei at target

$^{78}\text{Ni}$  :  $2 \times 10^9$  pps;

$^{132}\text{Sn}$  :  $2 \times 10^{11}$  pps

$^{91}\text{Kr}$ :  $1 \times 10^{12}$  pps;

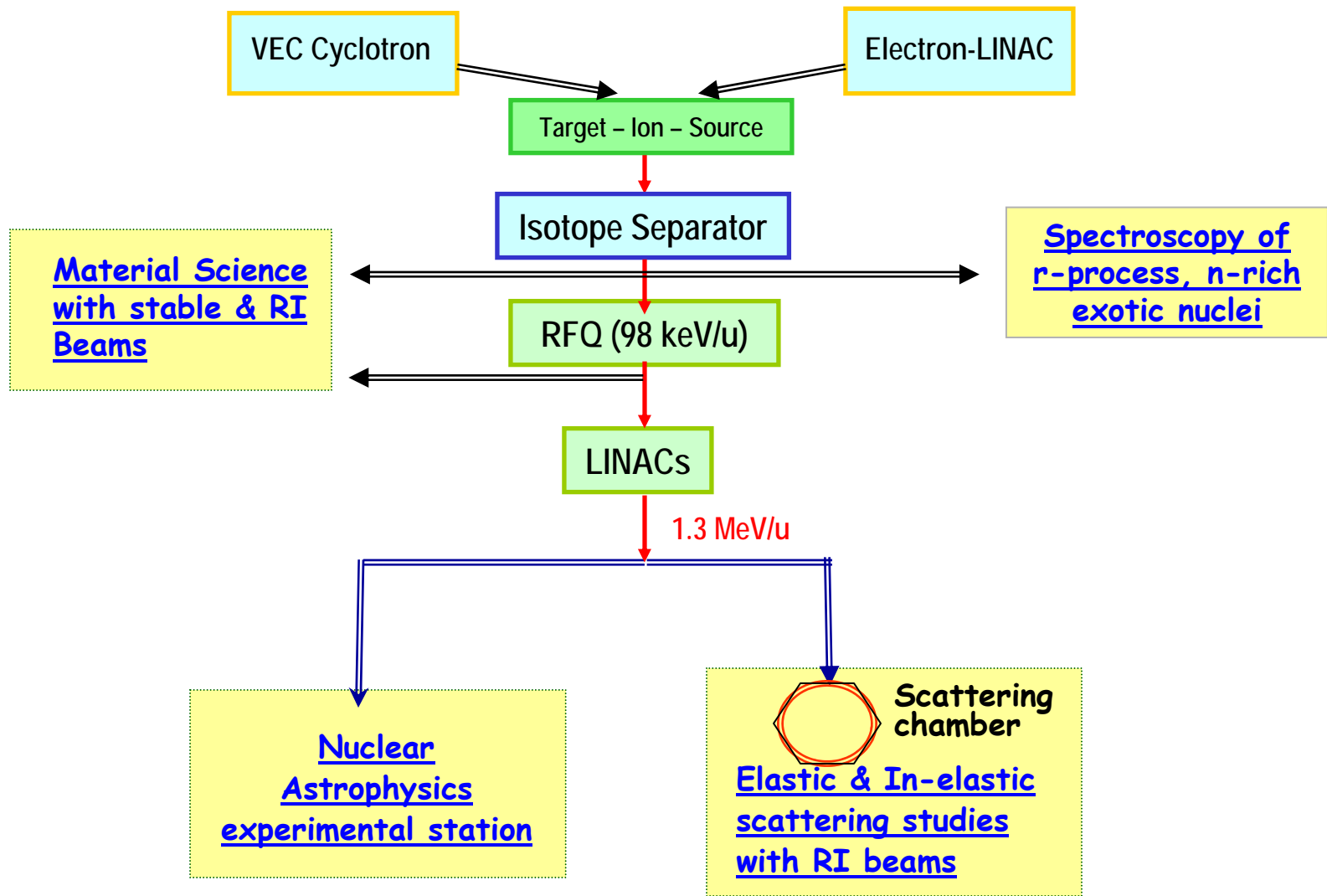
$^{94}\text{Kr}$ :  $3 \times 10^{10}$  pps

## Recent publication from RIB project group (in international peer review journals)

1. Phys. Lett. (in press), 2007. *Experiments*
2. Nucl. Instrum. & Meth. B261(2007)1018. *RIB facility status*
3. Rev Sci Instrum. Vol78 (2007) 043303. *RFQ results*
4. J of Phys. Condensed Matter 19, (2007) 236210. *Experiments*
5. Ceramics International, (in press). *Target*
6. Nucl. Instrum. & Meth. VolA560 (2006)182. *Linac design*
7. Nucl. Instrum. & Meth. VolA562 (2006)41. *Beam-line*
8. Nucl. Instrum. & Meth. VolA539 (2005)54. *Target*
9. Nucl. Instrum. & Meth. VolA547 (2005)270. *Charge breeder design*
10. J. of Mat. Sc. 40 (2005) 5265. *Experiments*
11. Nucl. Instrum. & Meth. VolA535 (2004)599. *RFQ design*
12. Physica C, Vol416, (2004) 25. *Experiments*
13. Nanotechnology 15 (2004) 1792. *Target*
14. Nucl. Instrum. & Meth. VolA447 (2000)345. *Charge breeder design*



Thank You!



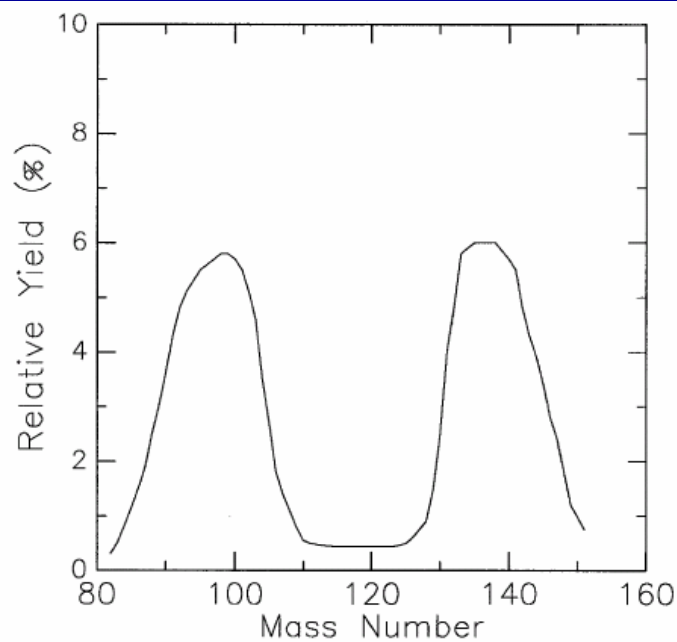
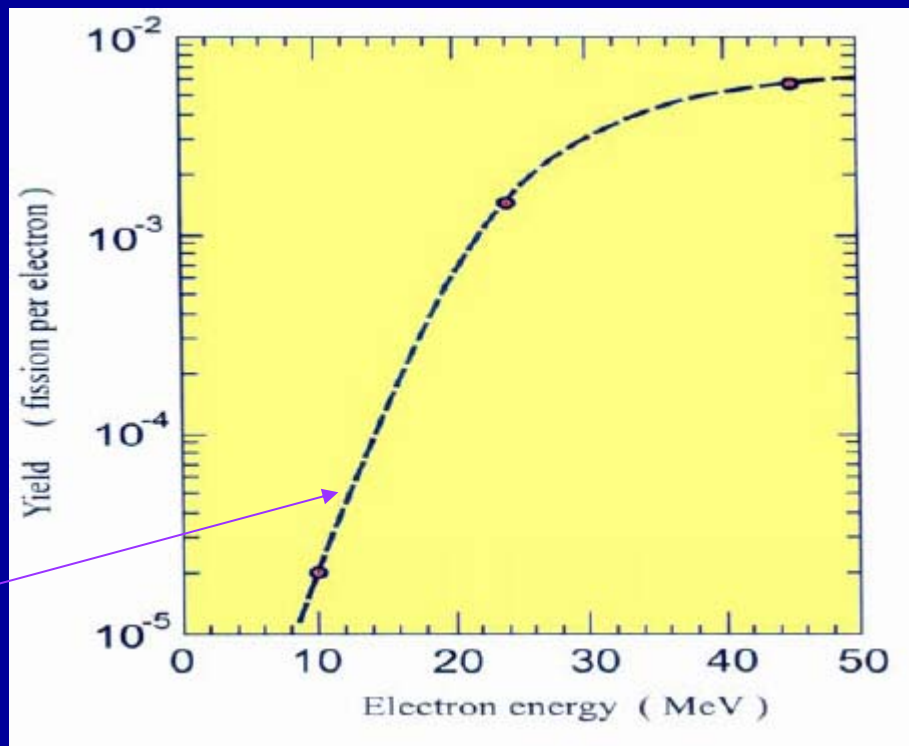


Fig. 10. The mass yield of fission fragments from electron-induced fission of  $^{238}\text{U}$  produced by 30 MeV electrons [32].

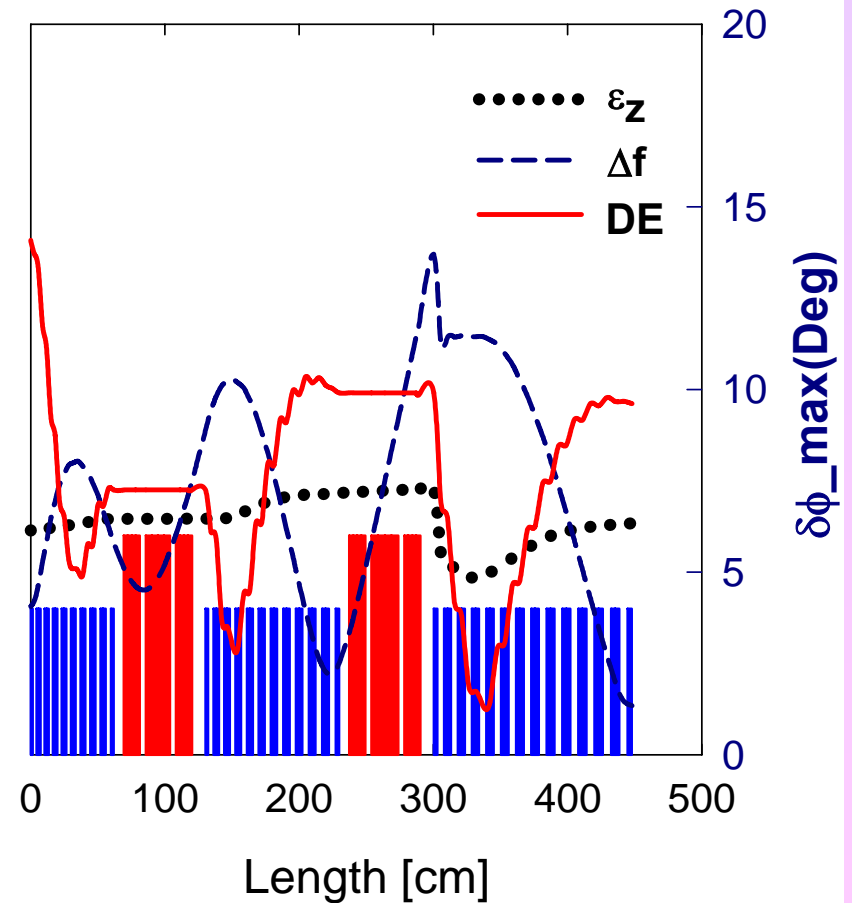
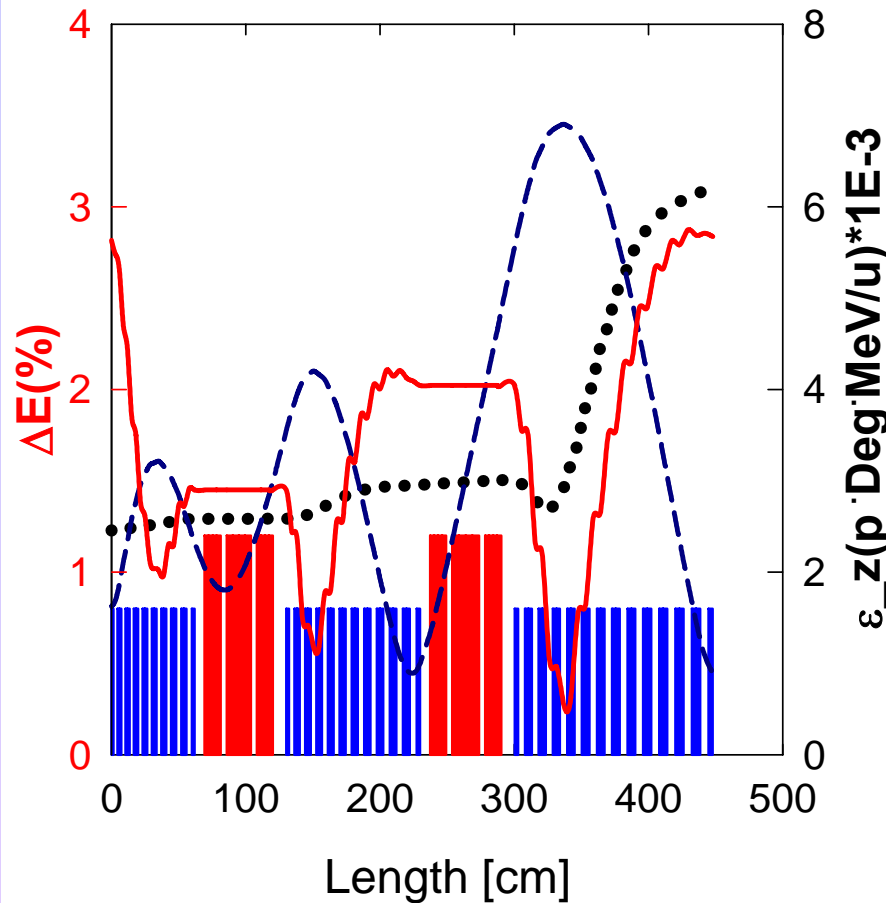
238U photo-fission fragments mass distribution

fission yield per electron for 238U as a function of electron energy



Better Intensity (95% transmission)

Better Beam Quality



Better Intensity (95% transmission)

$\epsilon_z$ -rms : increased by 2.5 times

$\Delta E$  : No change

$Df$  :  $4.06^\circ \rightarrow 2.23^\circ$

Better time structure (66% transmission)

No change

$2.8 \rightarrow 1.9\%$  (2\*rms energy width)

$4.06^\circ \rightarrow 1.34^\circ$  (2\*rms phase width)

# IH - LINAC : The design details →

## Energy Tunability

Possible set of energy tunes

Knobs : RF voltage & phase

Energy (keV/u)	Tank-1	Tank-2	Tank-3
263 – 397.5	Design	Design	Knobs
158.2 – 263	Design	Knobs	Buncher
100 – 158.2	Knobs	Off	Buncher

Captures >90% of beam

$$\left\{ \begin{array}{l} 2*\Delta\Phi_{rms} \\ 2*(\Delta T/T)_{rms} \end{array} \right.$$

Typ. Energy widths ~ ± 0.5-1%

Typ. time widths ~ ± 1 nS

Voltage & Phase settings

